

# Measurement of the neutrino velocity with the OPERA detector in the CNGS beam

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on behalf of the **OPERA** collaboration



**bmb+f** - Förderschwerpunkt  
**OPERA**  
Großgeräte der physikalischen  
Grundlagenforschung

# the OPERA collaboration

~160 scientists, 30 institutes, 11 countries

**Belgium**  
IIHE-ULB Brussels



**Croatia**  
IRB Zagreb



**France**  
LAPP Annecy  
IPNL Lyon  
IPHC Strasbourg



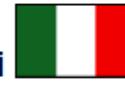
**Germany**  
Hamburg



**Israel**  
Technion Haifa



**Italy**  
LNGS Assergi  
Bari



Bologna  
LNF Frascati  
L'Aquila  
Naples  
Padova  
Rome  
Salerno

**Korea**  
Jinju



**Russia**  
INR RAS Moscow



LPI RAS Moscow  
ITEP Moscow  
SINP MSU Moscow  
JINR Dubna



**Japan**  
Aichi  
Toho  
Kobe  
Nagoya  
Utsunomiya



**Switzerland**  
Bern  
ETH Zurich



**Turkey**  
METU Ankara

additional contribution for neutrino velocity measurement:

**CERN**: CNGS, survey, timing and PS groups

**PTB** (National metrology institute, Germany)

**METAS** (National metrology institute, Switzerland)

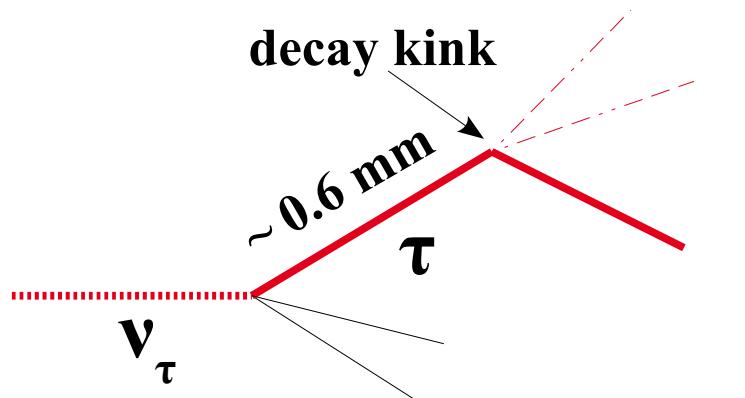
**Università Sapienza** (Rome University, Italy): Geodesy group

# overview

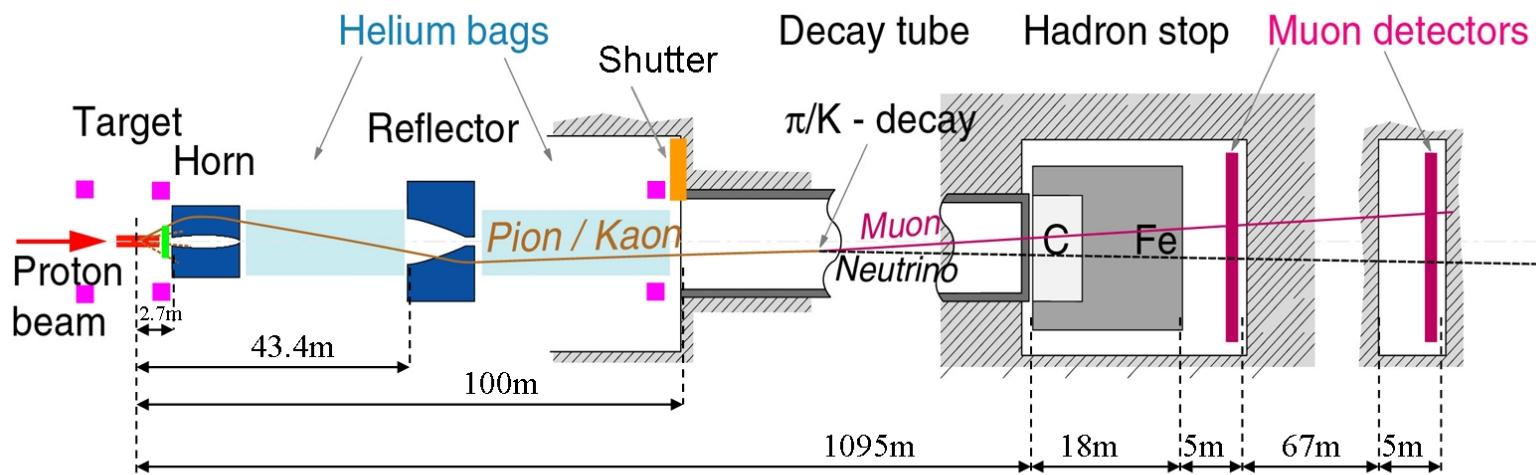
- the OPERA experiment
  - the CNGS neutrino beam
  - the OPERA detector
- neutrino time-of-flight measurement
  - experimental concept
  - 2011 results
  - recent developments
- summary

# the OPERA experiment

- main physics goal:
  - first direct detection of  $\nu_\mu \rightarrow \nu_\tau$  oscillations
- concept:
  - long baseline  $\nu_\mu$  beam,  $E_\nu \gg E_{\text{thresh}}(\text{CC } \nu_\tau) = 3.5 \text{ GeV}$
  - event-by-event detection of  $\tau$  leptons
- requirements:
  - high target mass ( $\sim 1000t$ )
  - high spatial resolution ( $\sim 1 \mu\text{m}$ )
  - very low background rate



# the CNGS neutrino beam

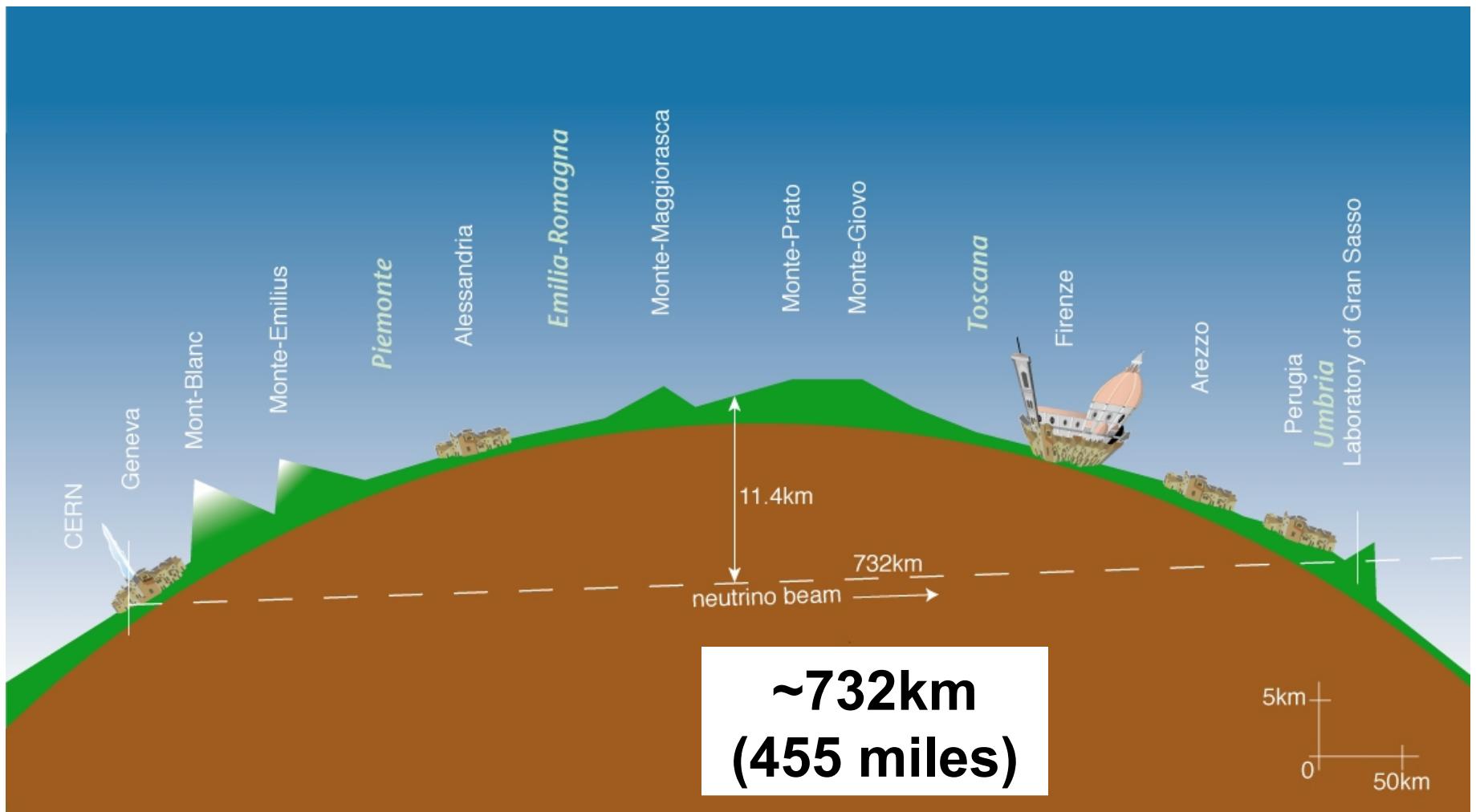


$$\langle E_{\nu_\mu} \rangle = 17 \text{ GeV}$$

$\bar{\nu}_\mu / \nu_\mu$	2.1%	(CC interactions)
$\nu_e / \nu_\mu$	0.89%	(CC interactions)
$\bar{\nu}_e / \nu_\mu$	0.06%	(CC interactions)
$\nu_\tau / \nu_\mu$	$< 10^{-4}\%$	(CC interactions)

about  $2.1 \times 10^{13}$  POT per extraction, 2 extractions per SPS filling

# neutrino propagation



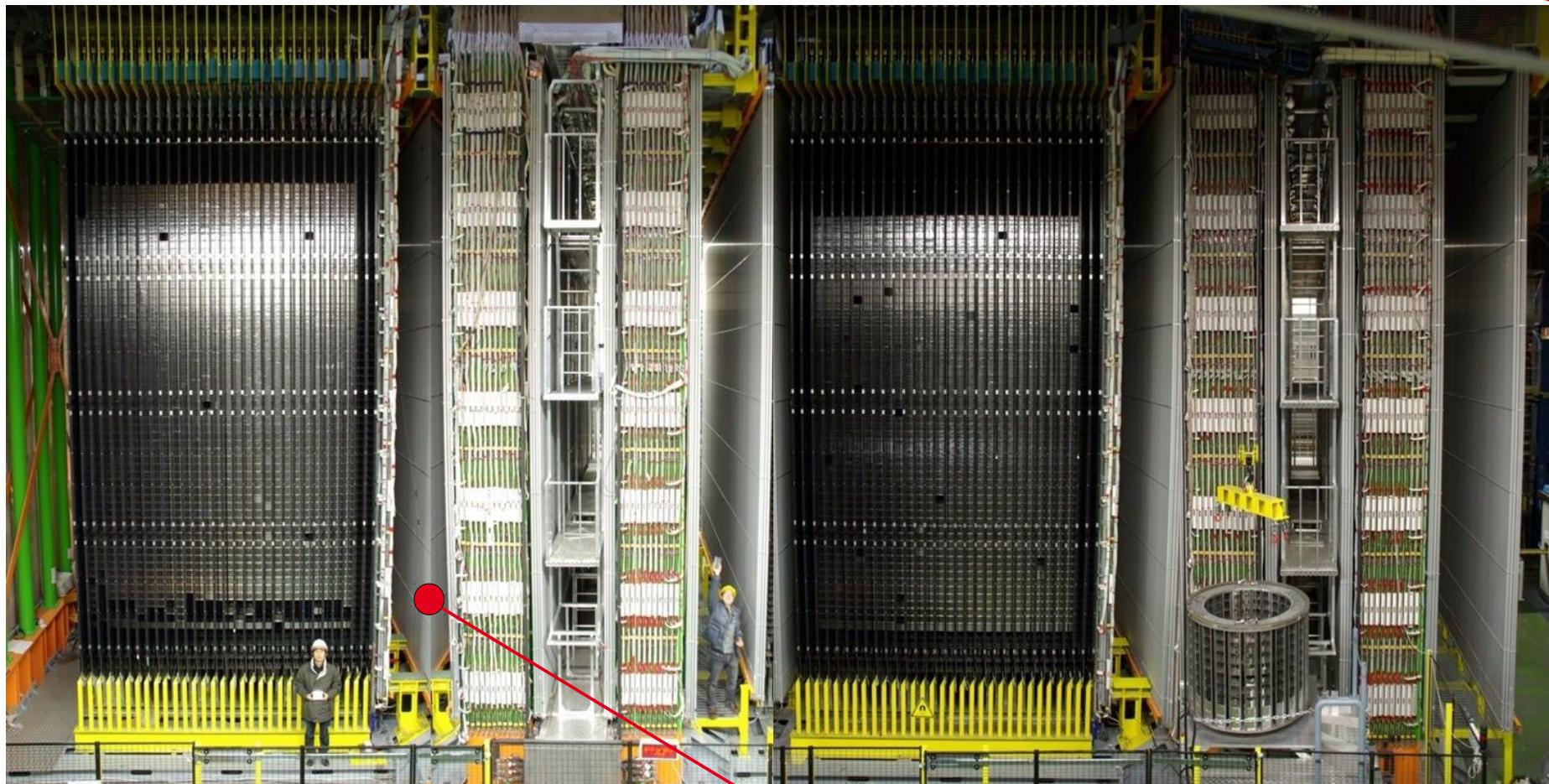
# neutrino propagation



- LNGS underground lab
  - under 1400m rock (3800mwe)
  - highway access



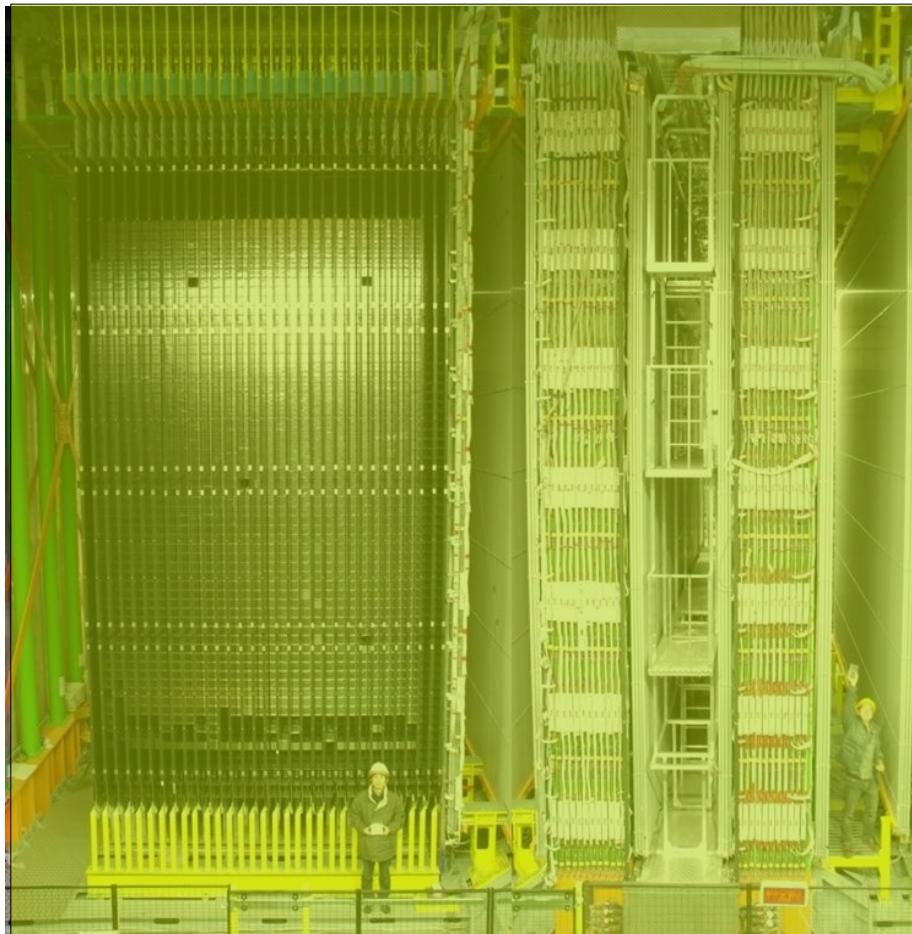
# the OPERA detector



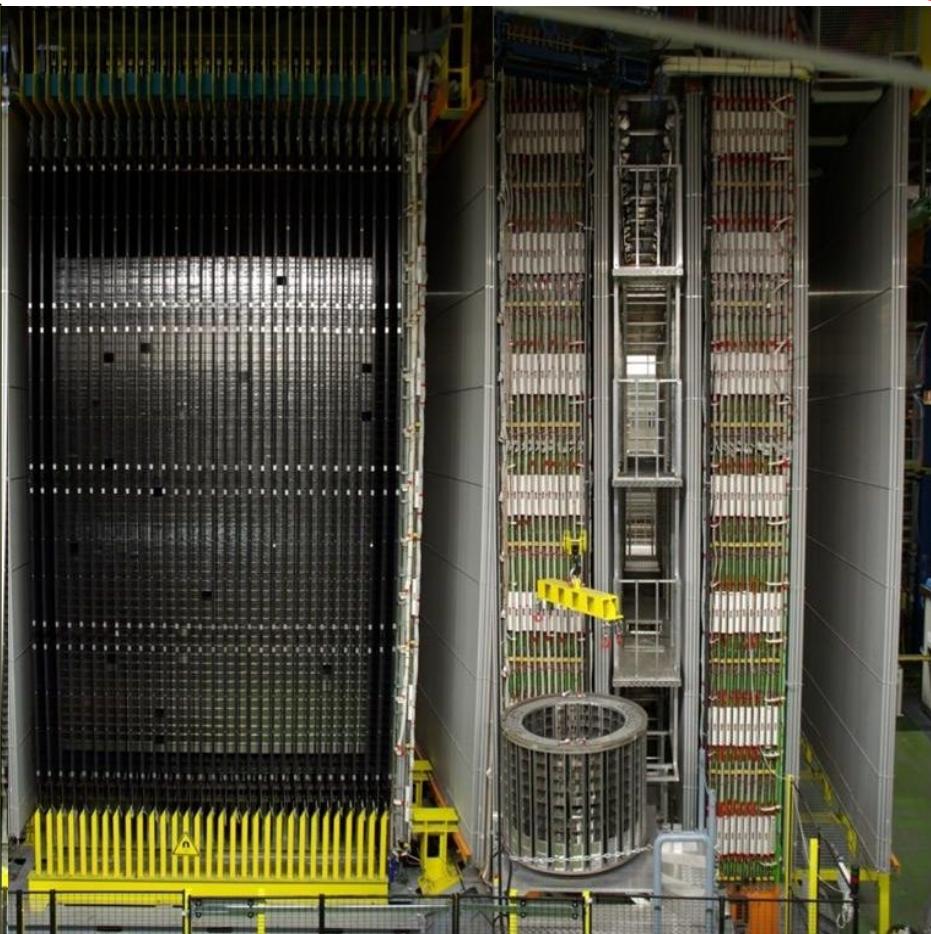
neutrinos →

reference point for TOF measurement: "A1"

# the OPERA detector

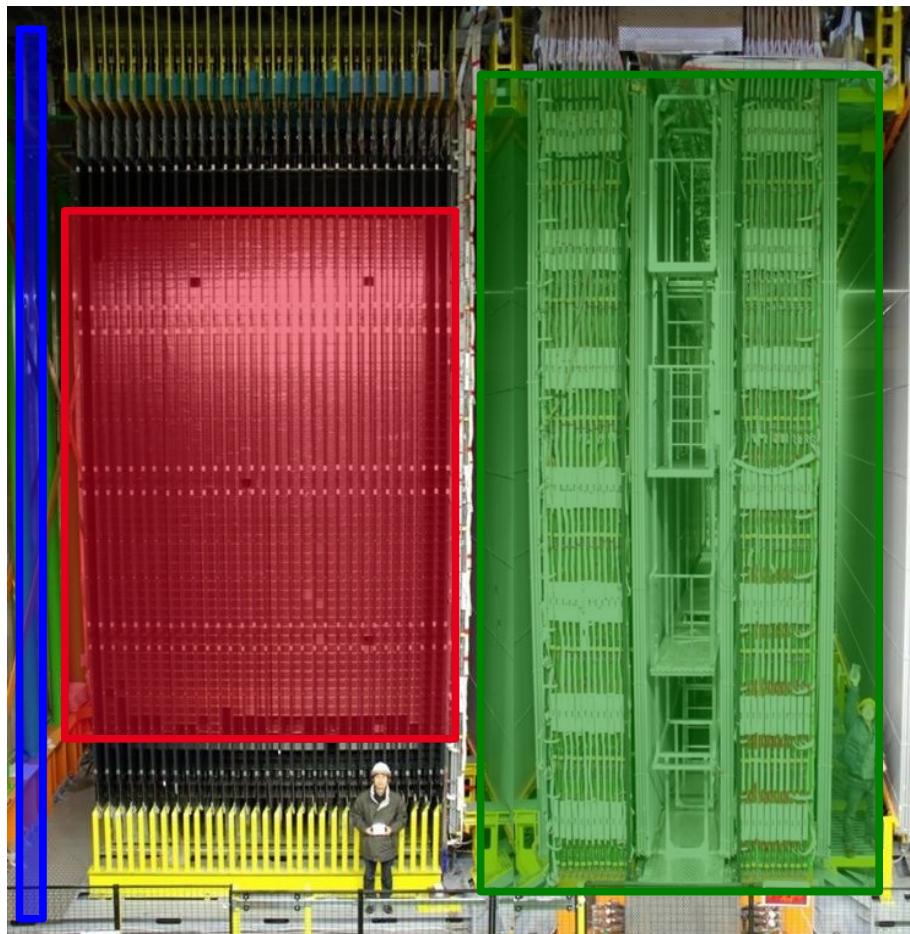


super module 1



super module 2

# the OPERA detector



super module 1

**veto (only SM1)**  
RPC

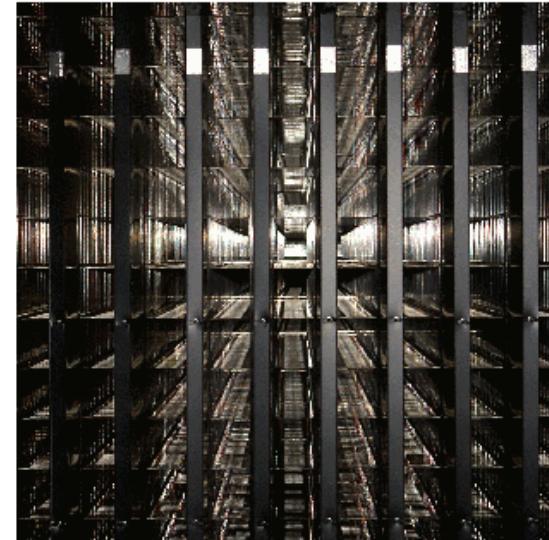
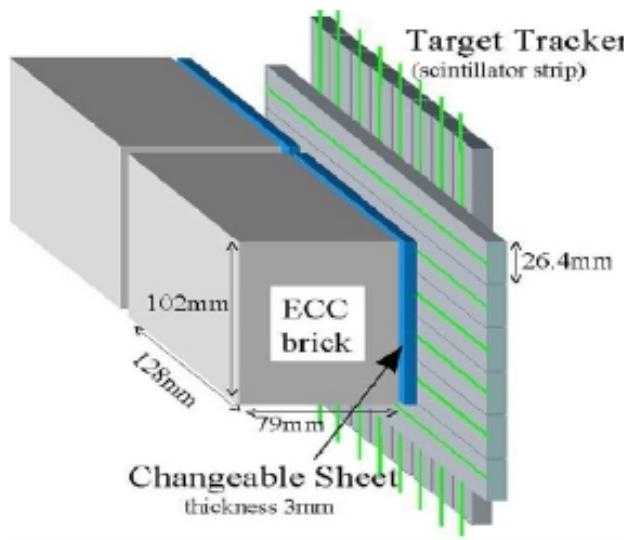
**target section**  
75,000 ECC bricks per SM  
31 pairs of planes of  
horiz. and vert. plastic  
scintillator strips

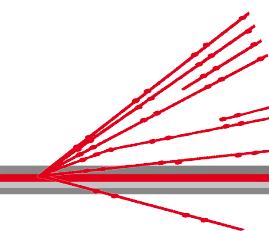
**spectrometer**  
1.5T dipole magnet  
RPC inner trackers  
drift tubes

# target section

## ■ 31 walls per SM

- lead/emulsion ECC
  - Changeable Sheets
  - horizontal scintillator strips
  - vertical scintillator strips
- } passive, excellent spatial/angular resolution
- } active, excellent time resolution (~1ns),  
spatial resolution ~1cm





# time-of-flight measurement

# introduction

## ■ definition of time-of-flight (TOF)

$$\text{TOF}_\nu = t_B - t_A - \text{delays}$$

## ■ “typical” TOF measurement principle

- measure the neutrino production time  $t_A$
- measure the distance between production and detection
- measure the neutrino detection time  $t_B$

## ■ definition of neutrino velocity:

$$v_\nu = \text{distance}/\text{TOF}_\nu$$

## ■ blind analysis (delays)

# brief history

## ■ 1979: FNAL (*Phys. Rev. Lett.* 43 (1979) 1361)

- short distance, 30 GeV  $\nu_\mu$ , comparison of  $\nu_\mu$  and  $\mu$  TOF
- $|\nu - c|/c \leq 4 \times 10^{-5}$

## ■ 1988: SN1987A (*Phys. Lett. B* 201 (1988) 353)

- very long distance (168,000 light years), 10 MeV anti- $\nu_e$ , comparison of  $\nu$  and photon arrival time (not SN mod.-dep.)
- $|\nu - c|/c \leq 2 \times 10^{-9}$

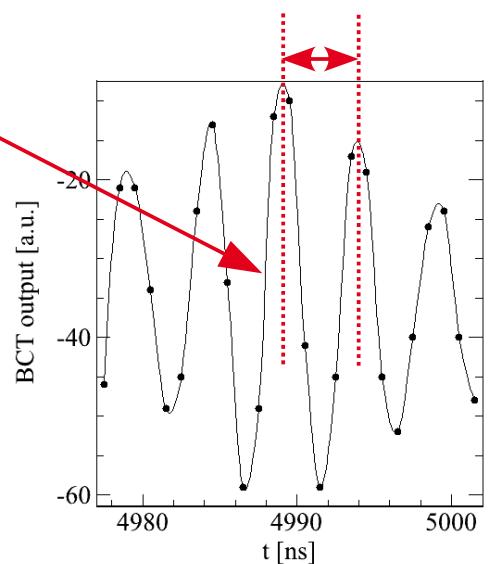
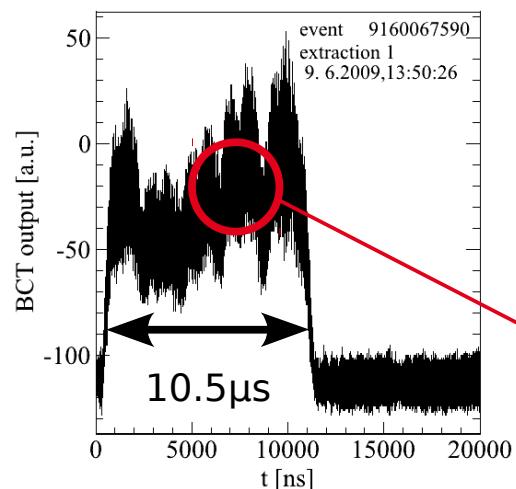
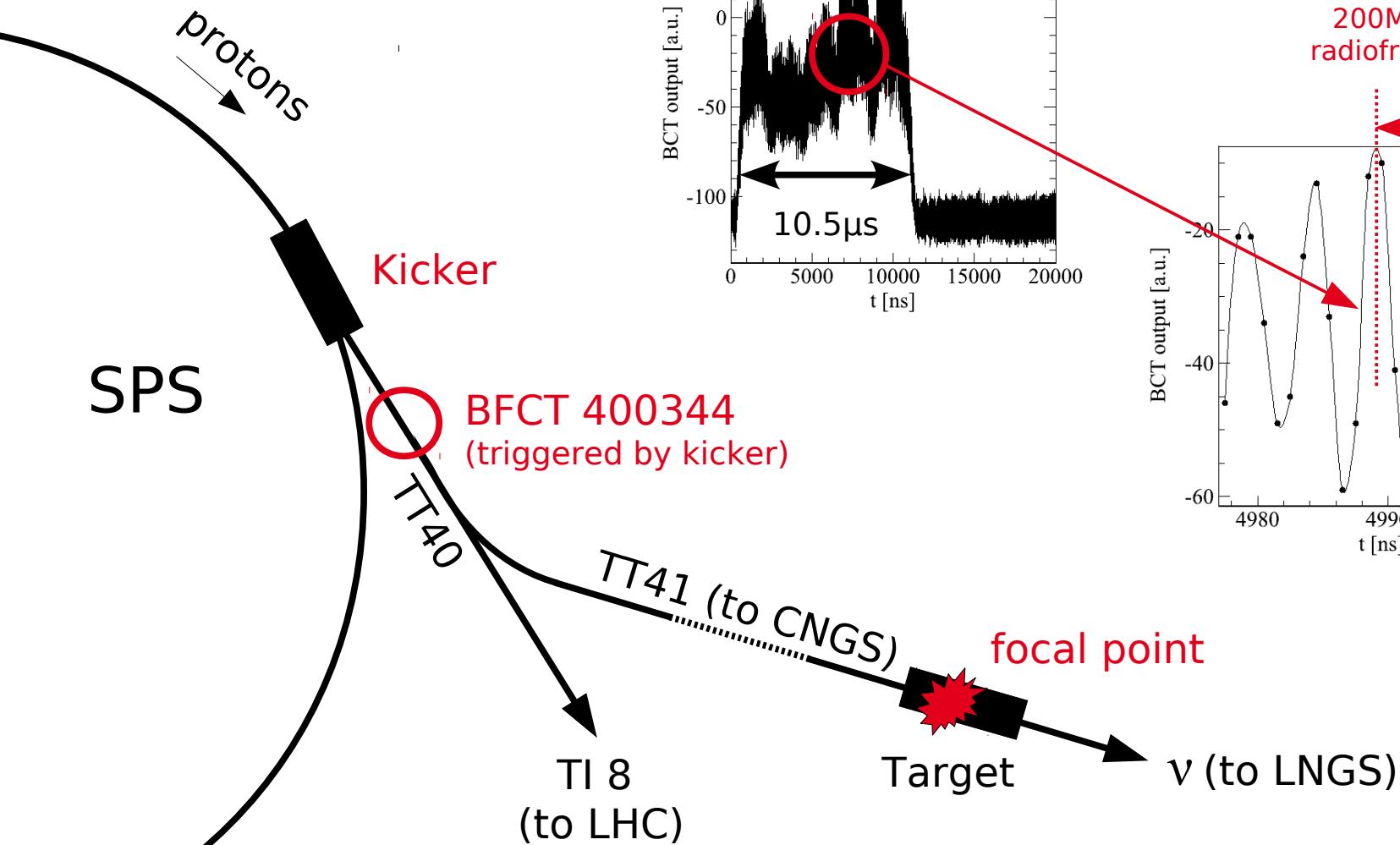
## ■ 2007: MINOS (*Phys. Rev. D* 76 (2007) 072005)

- 730km distance, ~3 GeV  $\nu_\mu$ , near detector comparison
- $(\nu - c)/c = (5.1 \pm 2.9) \times 10^{-5}$

## ■ 2011: OPERA

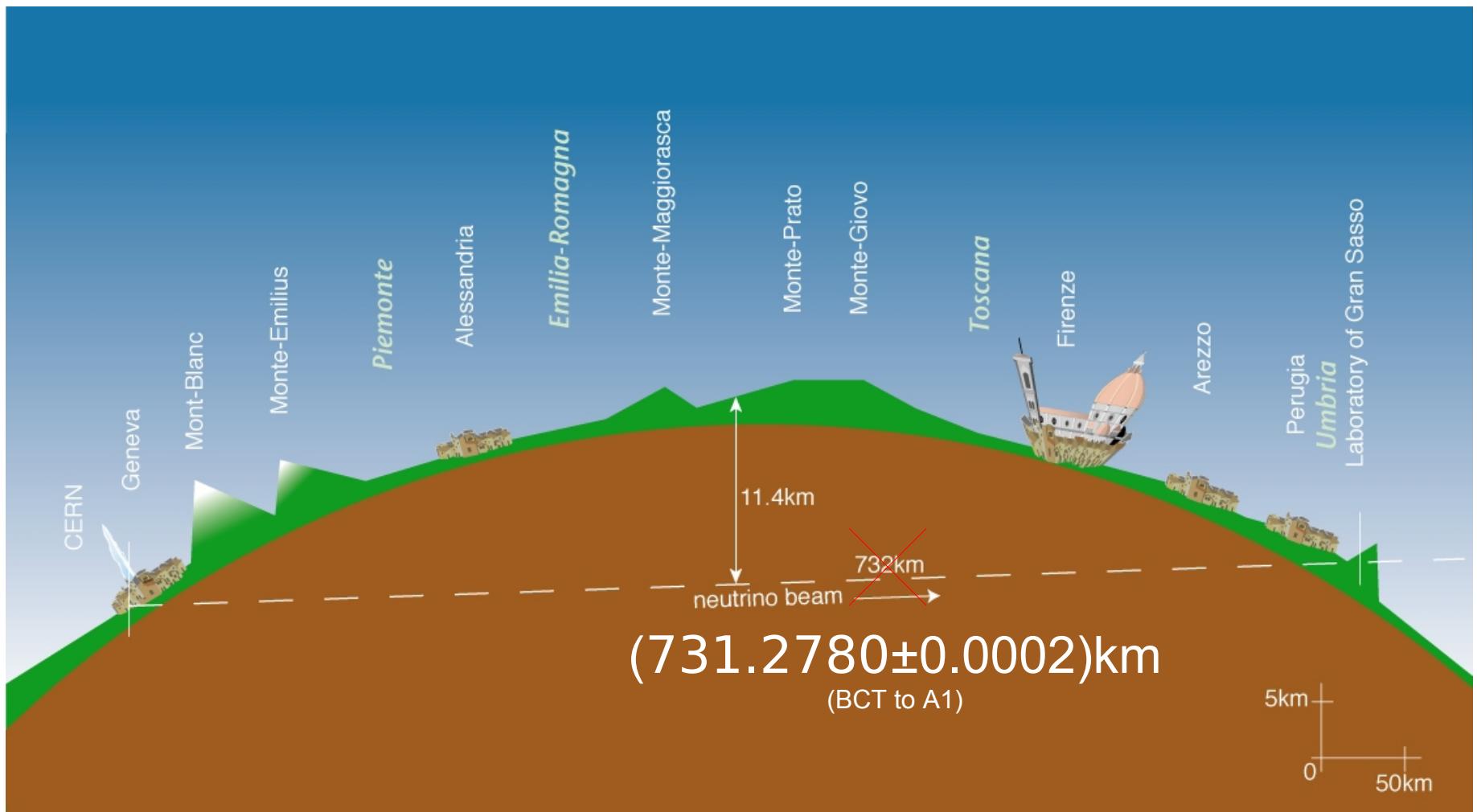
- 730km distance, ~17 GeV  $\nu_\mu$ , proton BCT comparison

# production time $t_A$

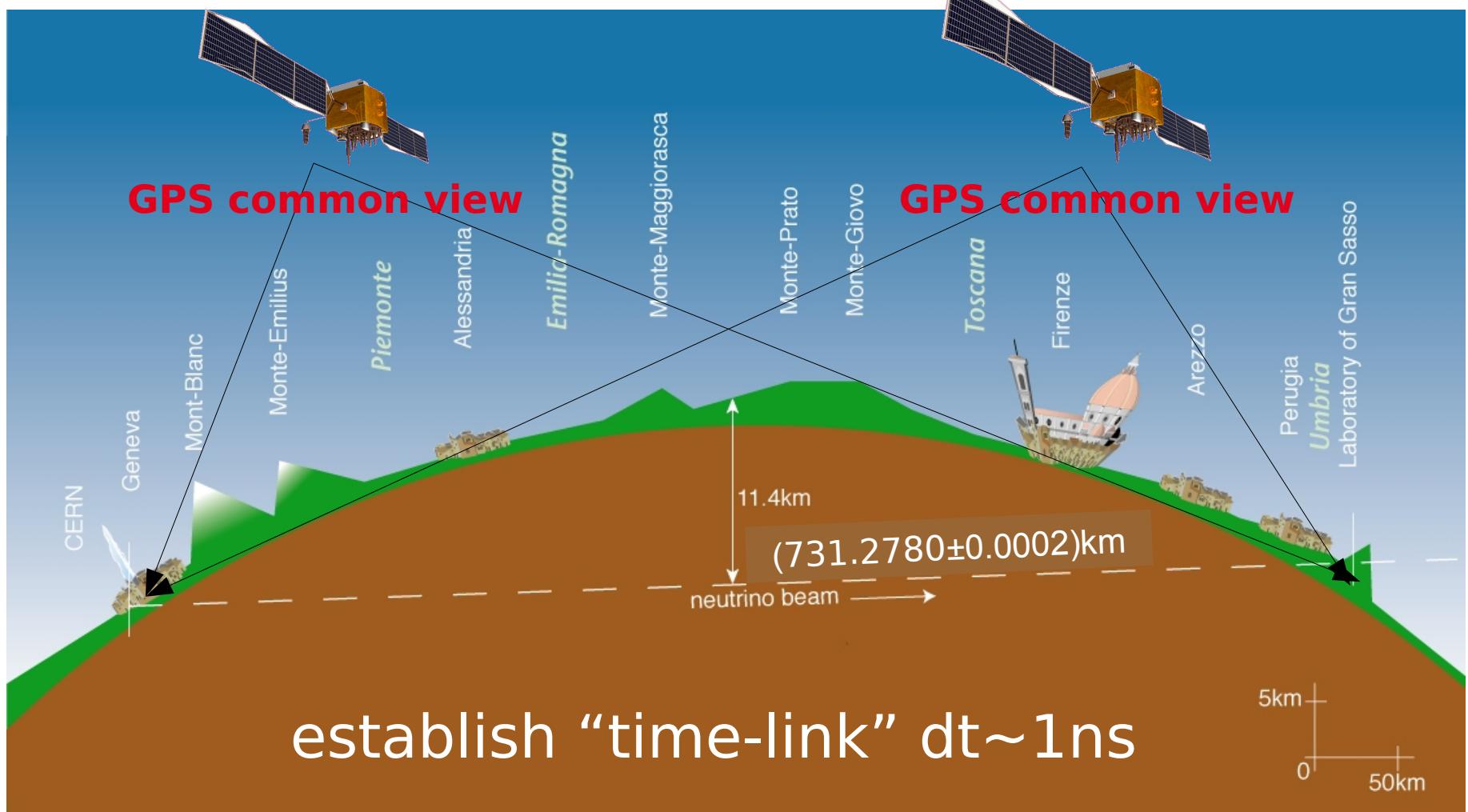


5ns spacing  
=  
200MHz SPS  
radiofrequency

# distance measurement

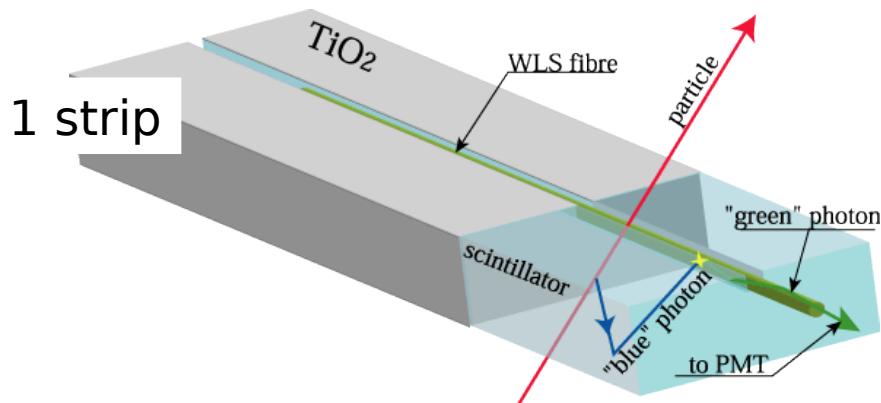
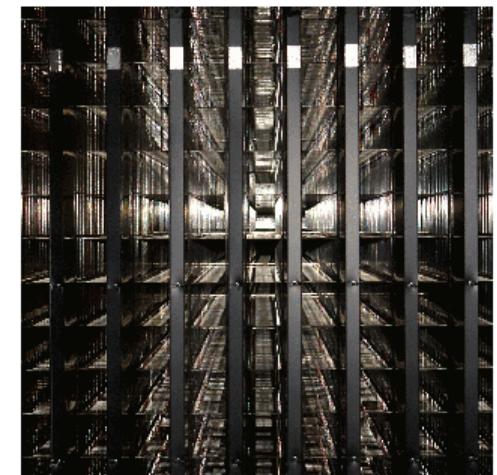
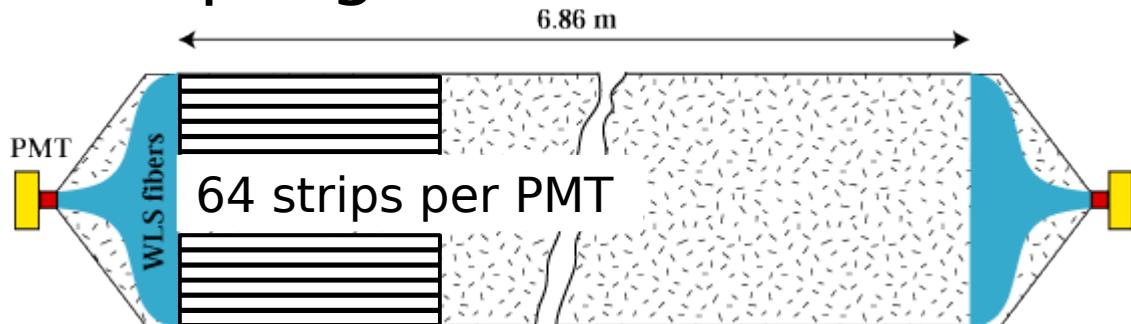


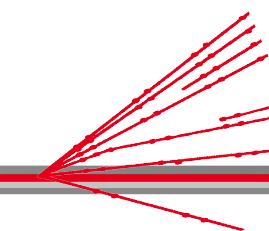
# clock synchronization



# neutrino detection time $t_B$

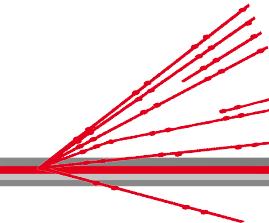
- use plastic scintillators only
- first hit in target trackers is the stop signal



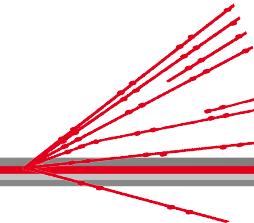


*September 2011*

# event selection



- selection of neutrino events
  - internal events (within fiducial volume, same as for the oscillation search): 7586
  - external events (interactions in rock) with reconstructed 3D muon track: 8525 ( $\pm 2\text{ns}$  additional uncertainty)
  - at least 4 satellites in common view
  - first hit not isolated in time or space
- 7235 internal and 7988 external events
- if neutrino event passes selection:  
select the corresponding BCT waveform



- original method:  
build likelihood from summed waveforms

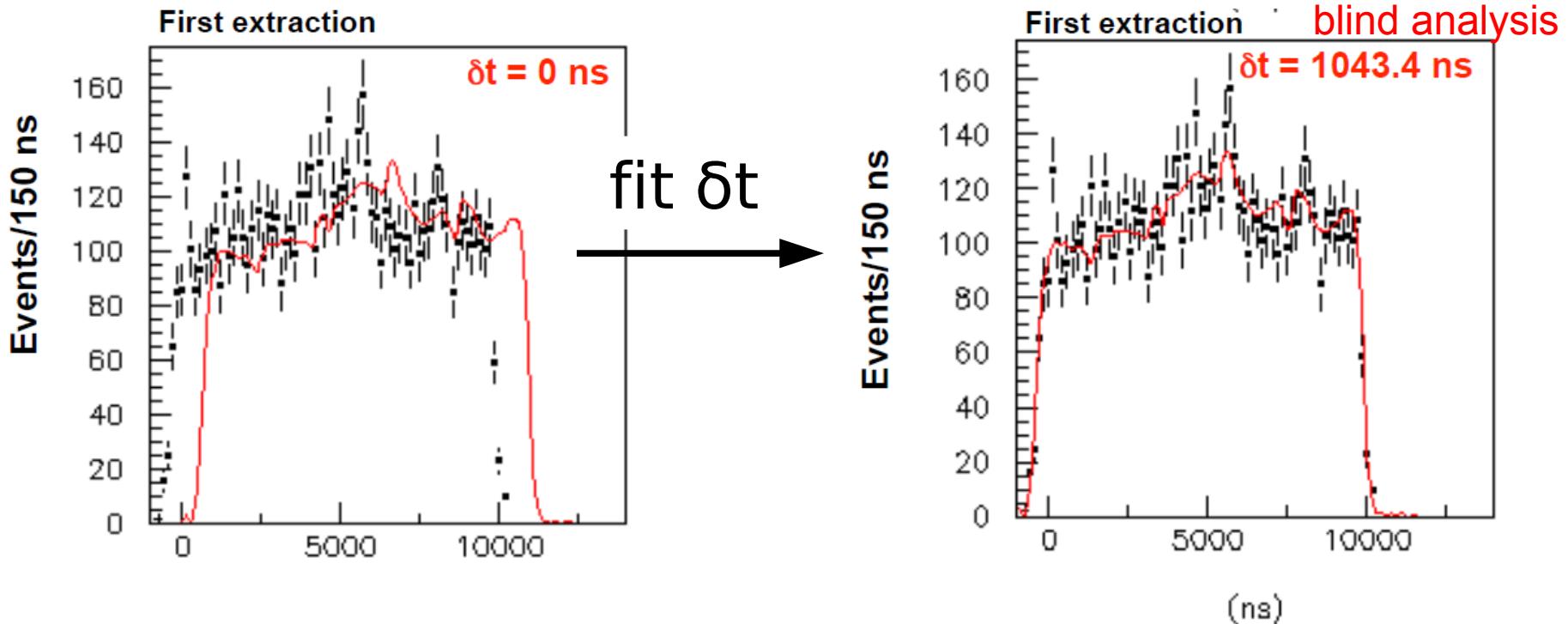
$$L_k(\delta t_k) = \prod_j w_k(t_j + \delta t_k) \quad k = 1, 2 \text{ extractions}$$

- alternative method:  
build likelihood from single waveforms,  
(smaller stat. uncertainty, additional syst.  
uncertainty):

$$L(\delta t) = \prod_j w_j(t_j + \delta t)$$

# analysis

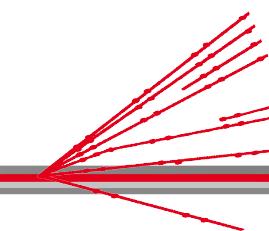
the red curve is for visualization only!



under investigation

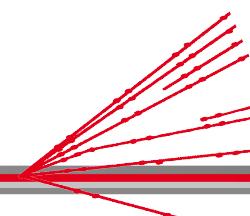
original method:  $(57.8 \pm 7.8 \text{ (stat.)} {}^{+8.3}_{-5.9} \text{ (sys.)}) \text{ ns}$

alternative method:  $(54.5 \pm 5.0 \text{ (stat.)} {}^{+9.6}_{-7.2} \text{ (sys.)}) \text{ ns}$



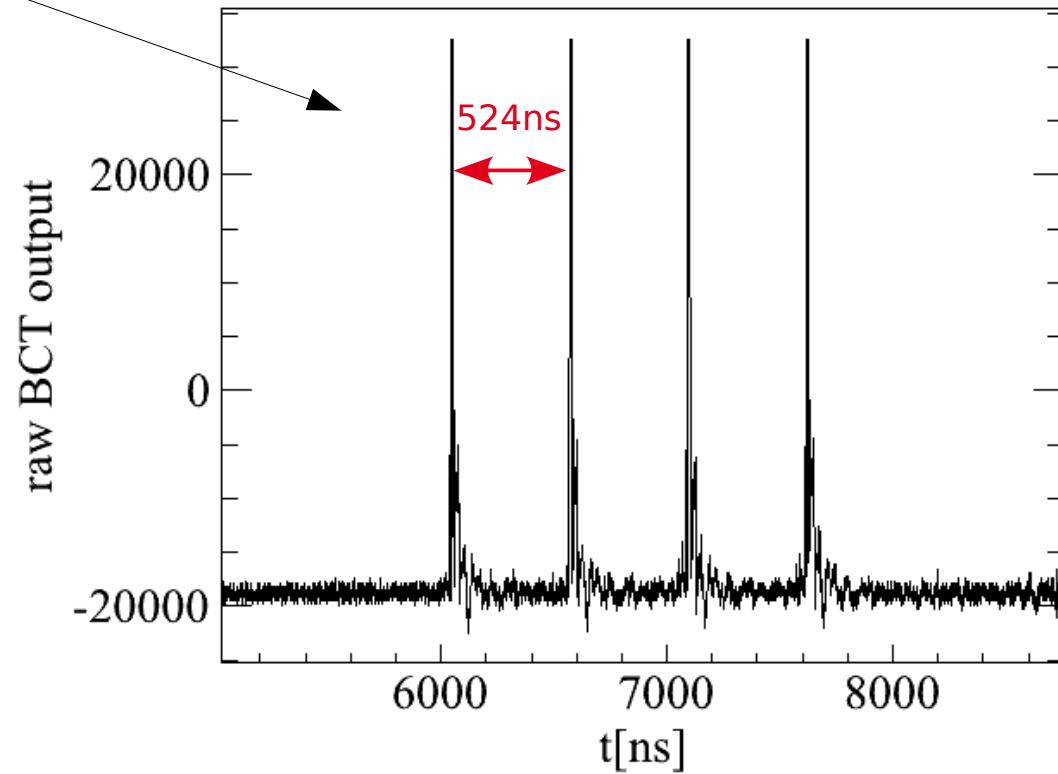
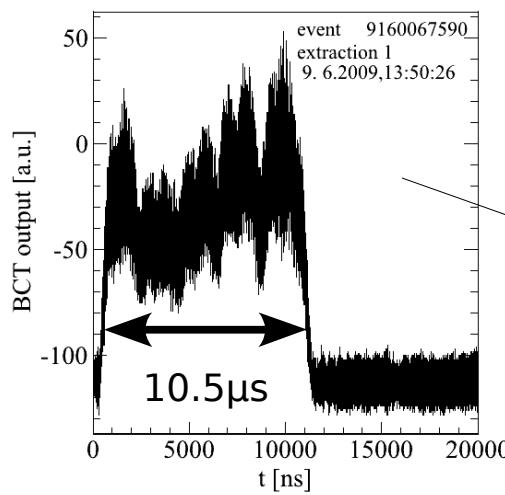
*November 2011*

# bunched beam (1)

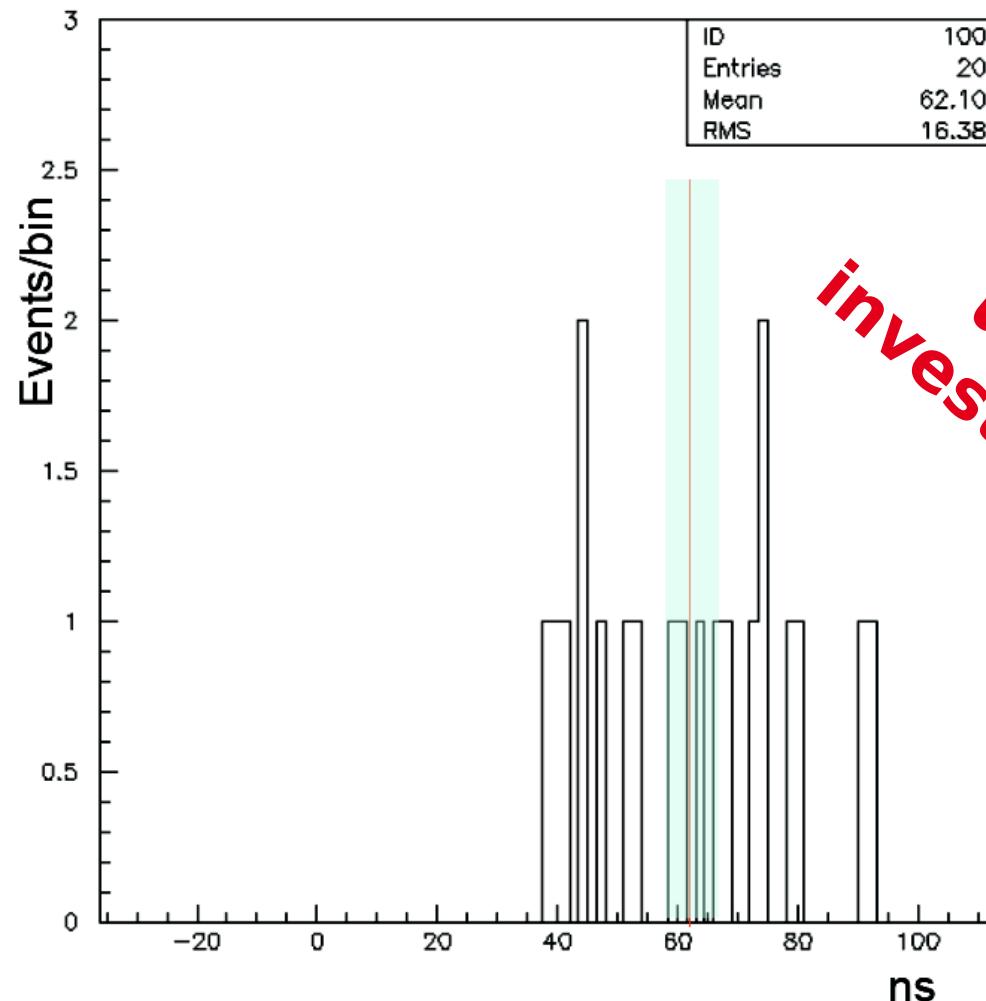


- bunched beam: instead of 10.5 $\mu$ s extractions:  
4 single, 3ns-wide bunches, separated by 524ns  
→ **single-event TOF measurement!**
  - October 22 to November 6, 2011
  - beam intensity lower than nominal ( $\sim 1/60$ )
  - collected 35 events, same selection criteria,  
same delay corrections
- **14 external and 6 internal events**

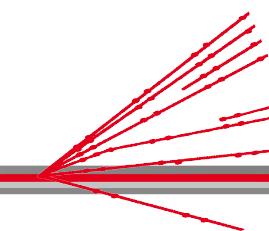
# bunched beam (2)



# bunched beam (3)

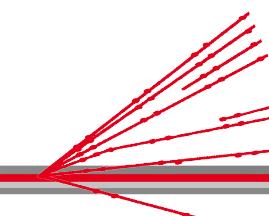


bunched beam method:  $(62.1 \pm 3.7 \text{ (stat.)} {}^{+8.3}_{-5.9} \text{ (sys.)}) \text{ ns}$



*December 2011 - February 2012*

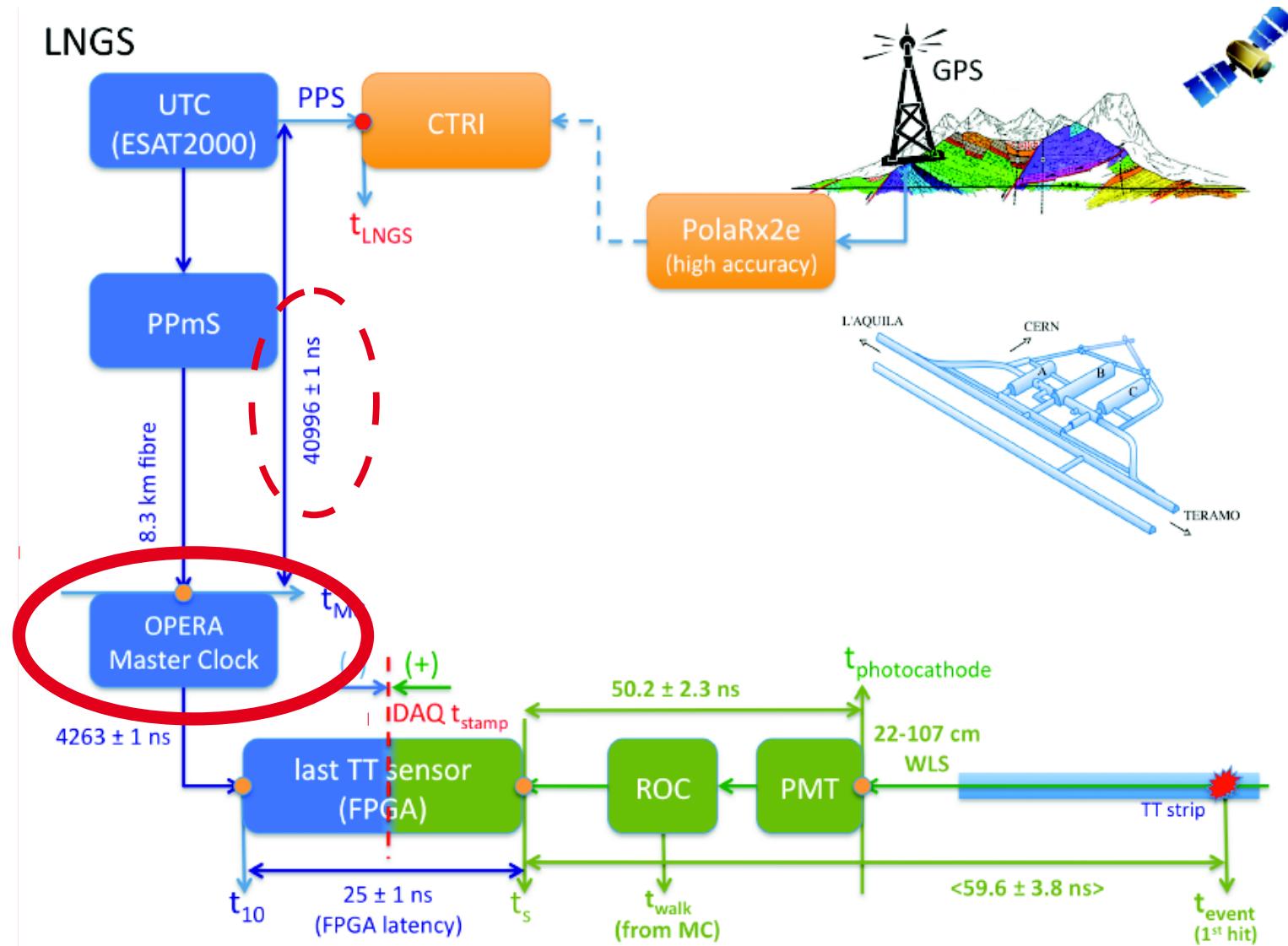
# OPERA statement

A red horizontal bar at the top right features a beam splitter icon: a red rectangle with a diagonal line and several red lines diverging from it to represent light particles.

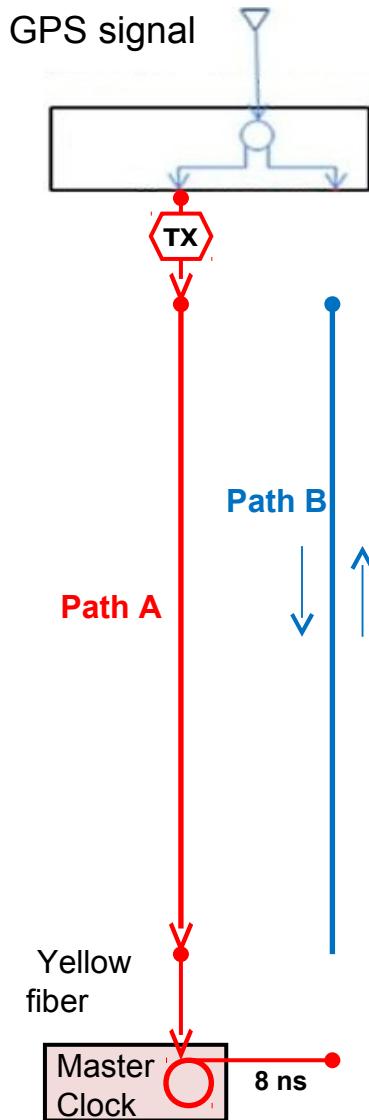
**“The OPERA Collaboration [...] has identified two issues that could significantly affect the reported result.**

[...] the **oscillator used to produce the events time-stamps [...]**  
[...] the **connection of the optical fiber [...]**” (Feb. 23<sup>rd</sup> 2012)

# LNGS timing

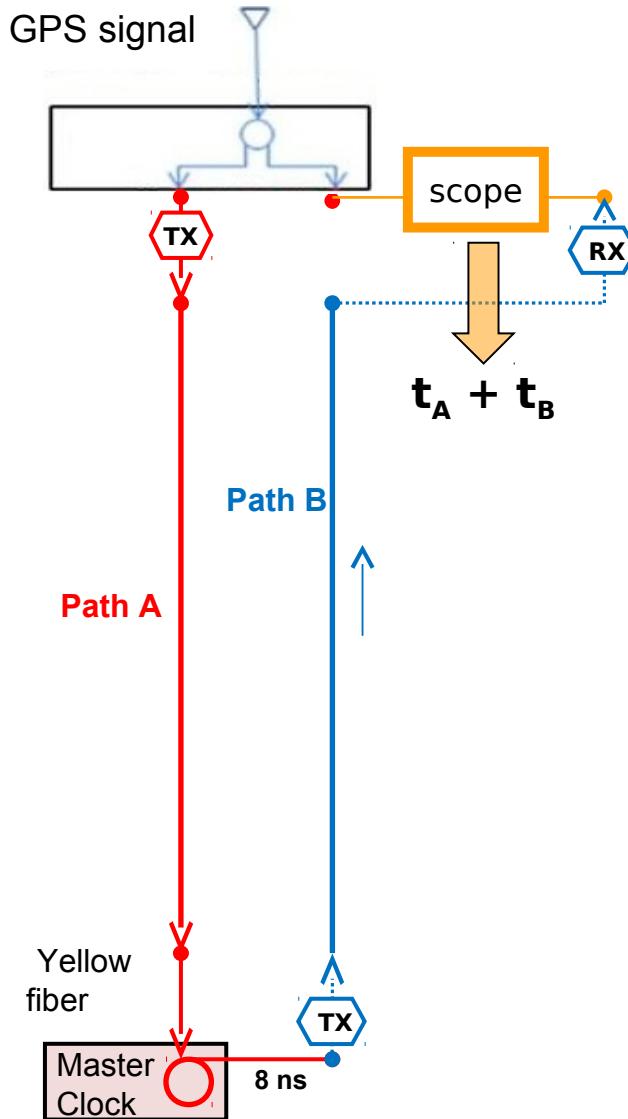


# connection of the optical fiber



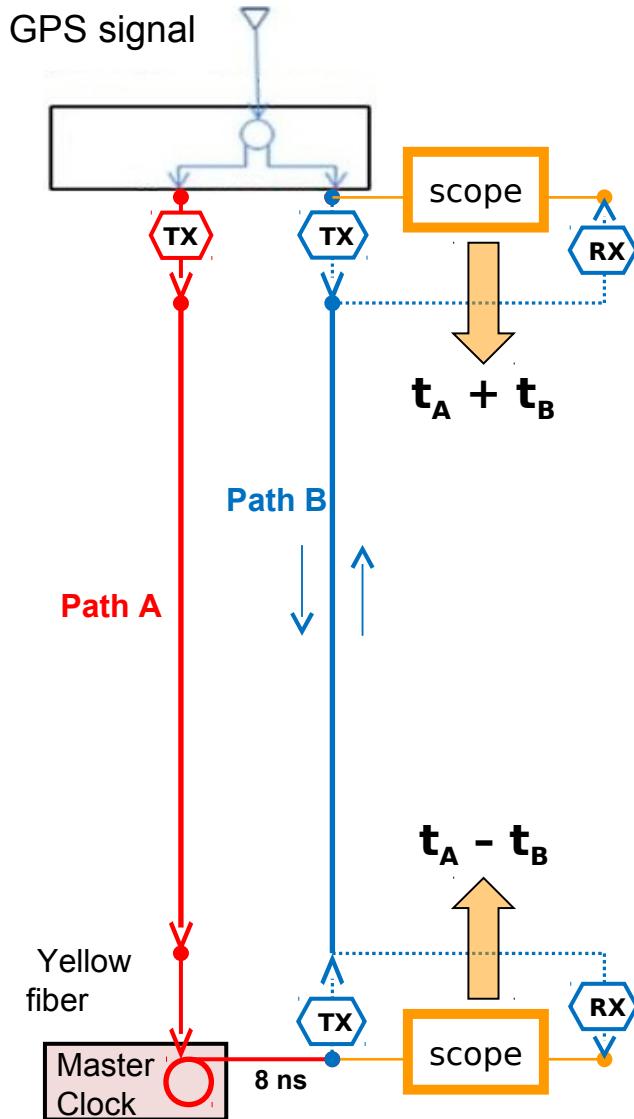
- ~8.3km long optical fiber
- dedicated campaign Dec11-Feb12
- “two ways measurements” using the same auxiliary fiber

# connection of the optical fiber



- ~8.3km long optical fiber
- dedicated campaign Dec11-Feb12
- “two ways measurements” using the same auxiliary fiber

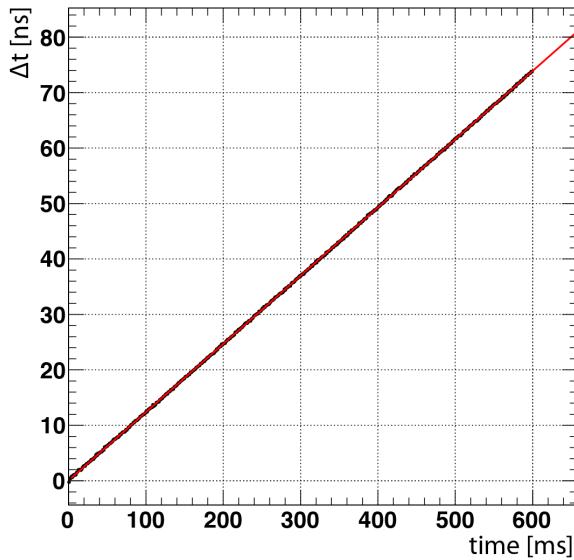
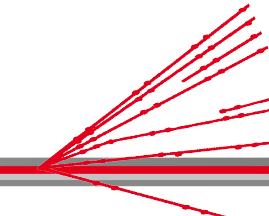
# connection of the optical fiber



- ~8.3km long optical fiber
- dedicated campaign Dec. 11-Feb. 12
- “two ways measurements” using the same auxiliary fiber
- identified issue: “yellow fiber” connection to Master Clock (MC), dependence of MC analogue circuit response to input light amplitude

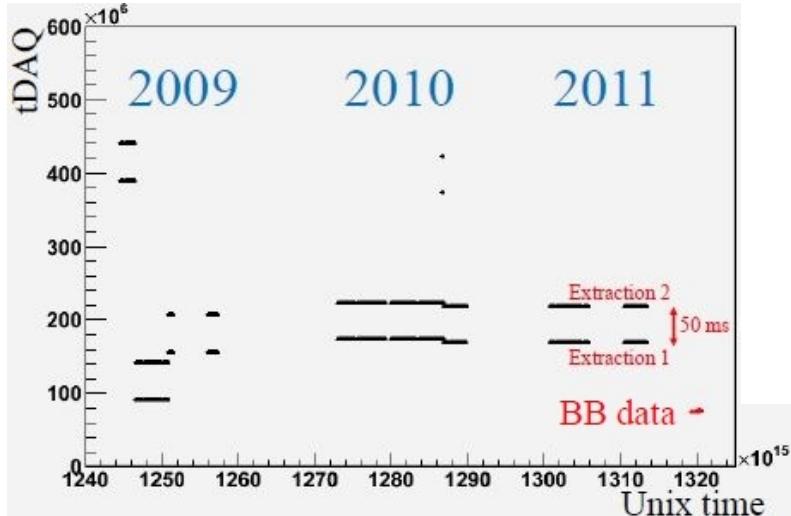
→ fiber delay was measured 74 ns larger in early Dec. 2011 than in previous measurements (before 2008), and in later ones (after 13 Dec. 2011): time when anomalous condition occurred and stability of it is under investigation

# oscillator (master clock)



- OPERA DAQ is reset every 0.6s
- Within 0.6s long cycle, DAQ timing is performed by an internal 20MHz counter of the Master Clock (MC)

→ internal MC frequency found too high by  $\Delta f/f = 1.23 \times 10^{-7}$ , OPERA DAQ timestamps are delayed, measured TOF is up to 74ns too long (depending on position in DAQ cycle)



→ CNGS events are NOT equally distributed within each DAQ cycle (event-by-event correction)

# TOF summary (1)

- OPERA TOF result under investigation
- time unkown when anomalous conditions occurred during data taking:

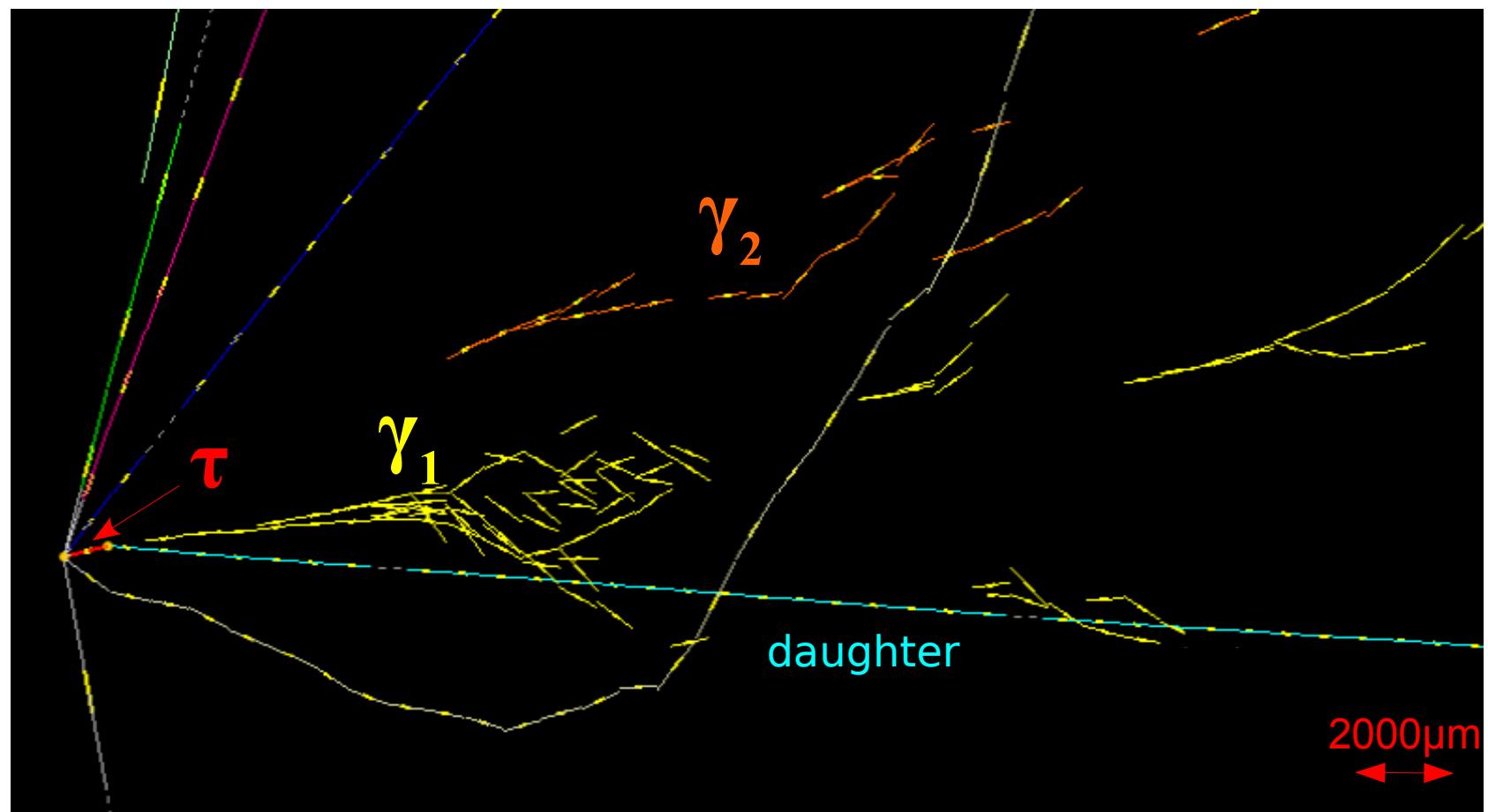
To add information, study of OPERA-LVD coincidences using cosmic muons is being finalized, release is foreseen **tomorrow** at a seminar at LNGS starting 14:30  
(<http://agenda.infn.it/conferenceDisplay.py?confId=4896>)

## TOF summary (2)

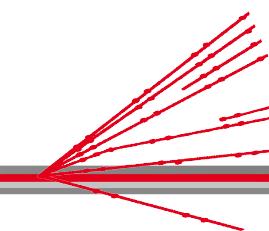
- While additional investigations are being performed to unambiguously quantify the size of the combined effects on the observed neutrino velocity result, the Collaboration is looking forward to perform a new measurement as soon as a new bunched beam will be available in 2012.
- Note: all LNGS experiments share the (BCT based) CERN timing, the GPS common view hardware and the CERN-LNGS distance measurement.

# overall summary

- OPERA is a neutrino oscillation experiment
- found 1  $\nu_\tau$  candidate, while  $0.05 \pm 0.01$  bkgd. events were expected. The analyzed sample corresponds to about 25% of the overall data collected until end of 2012
- electron neutrino appearance results in 2012
- no data taking in 2013 (CERN shutdown)

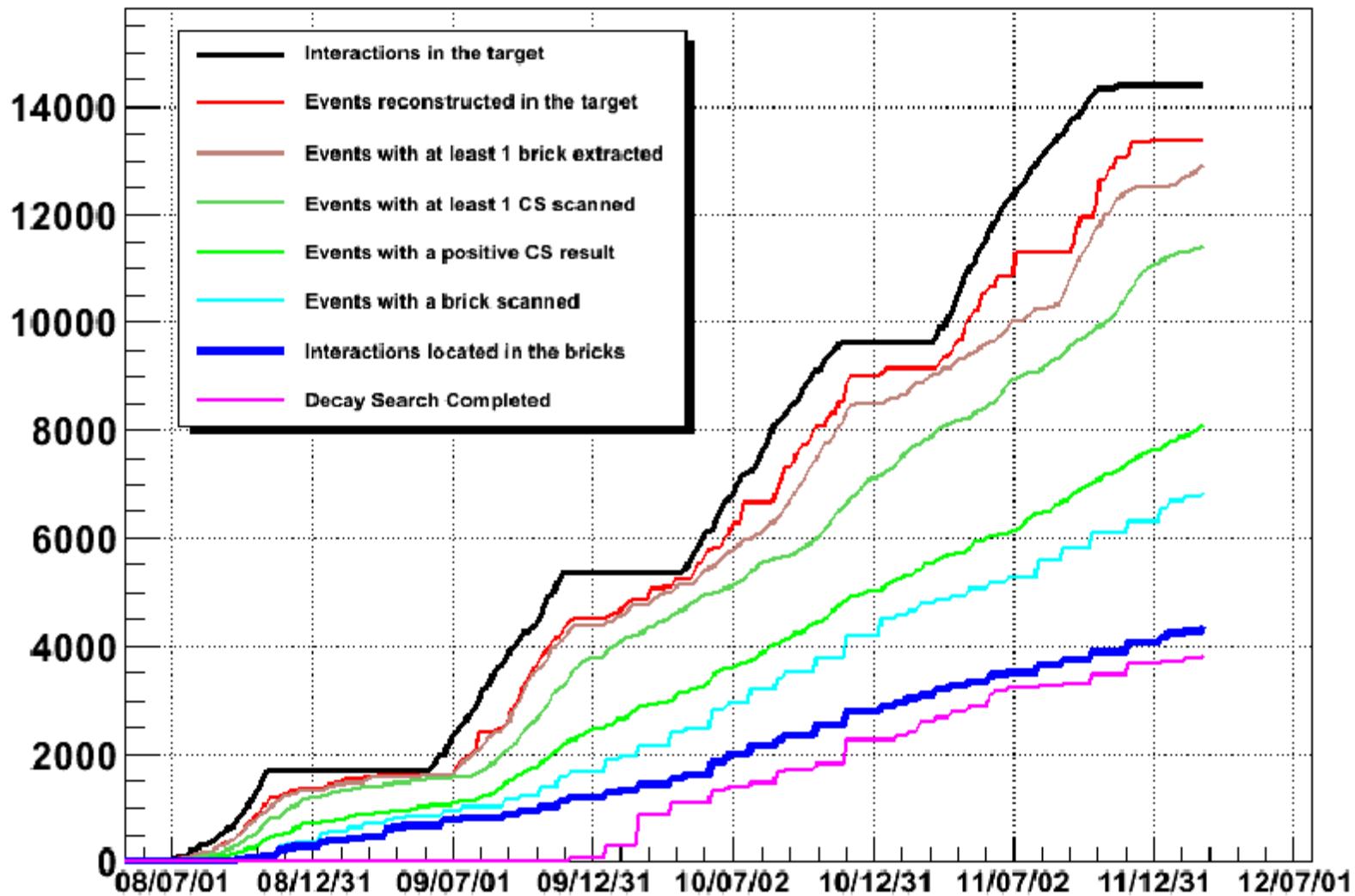


Thank you!

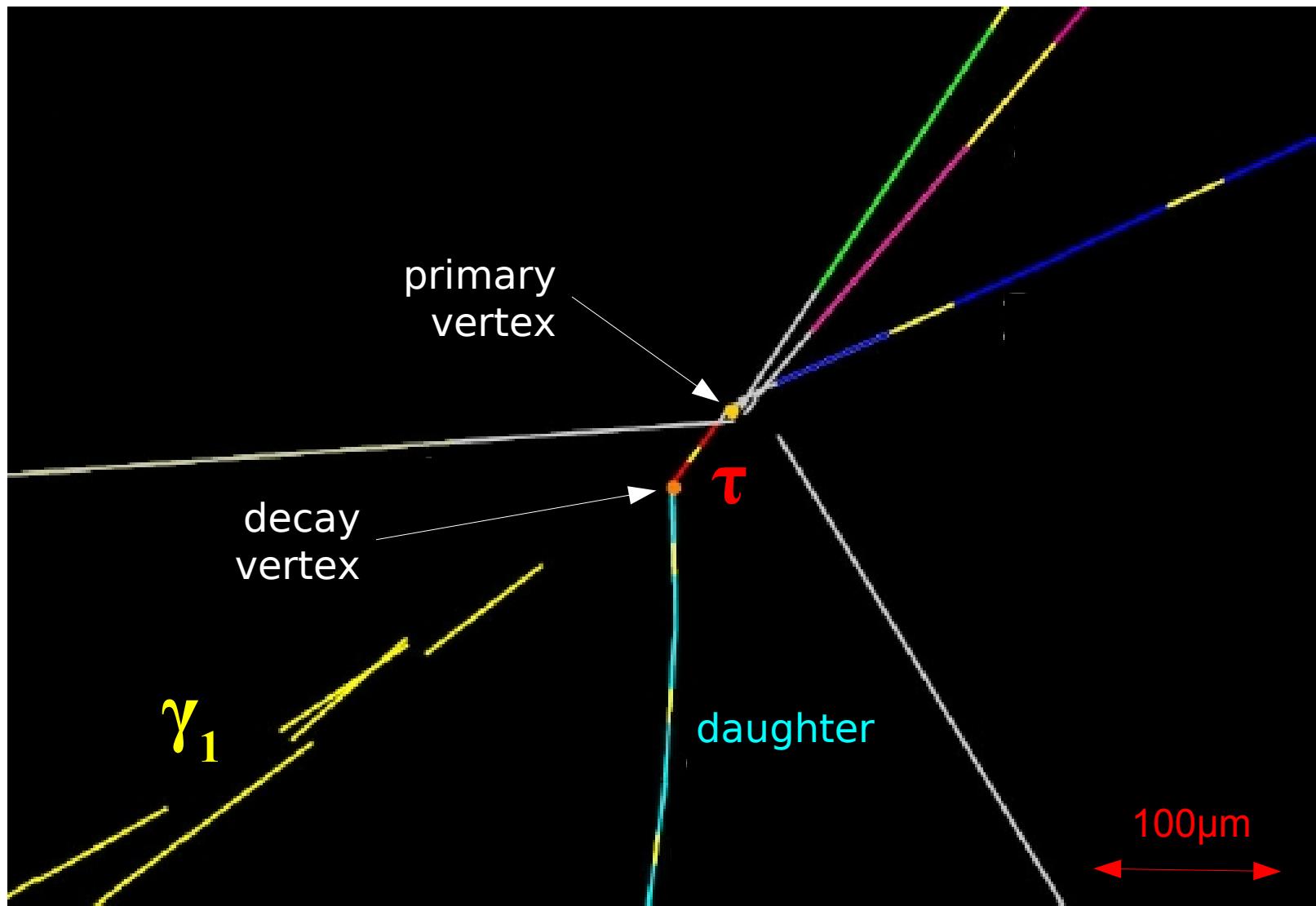


backup

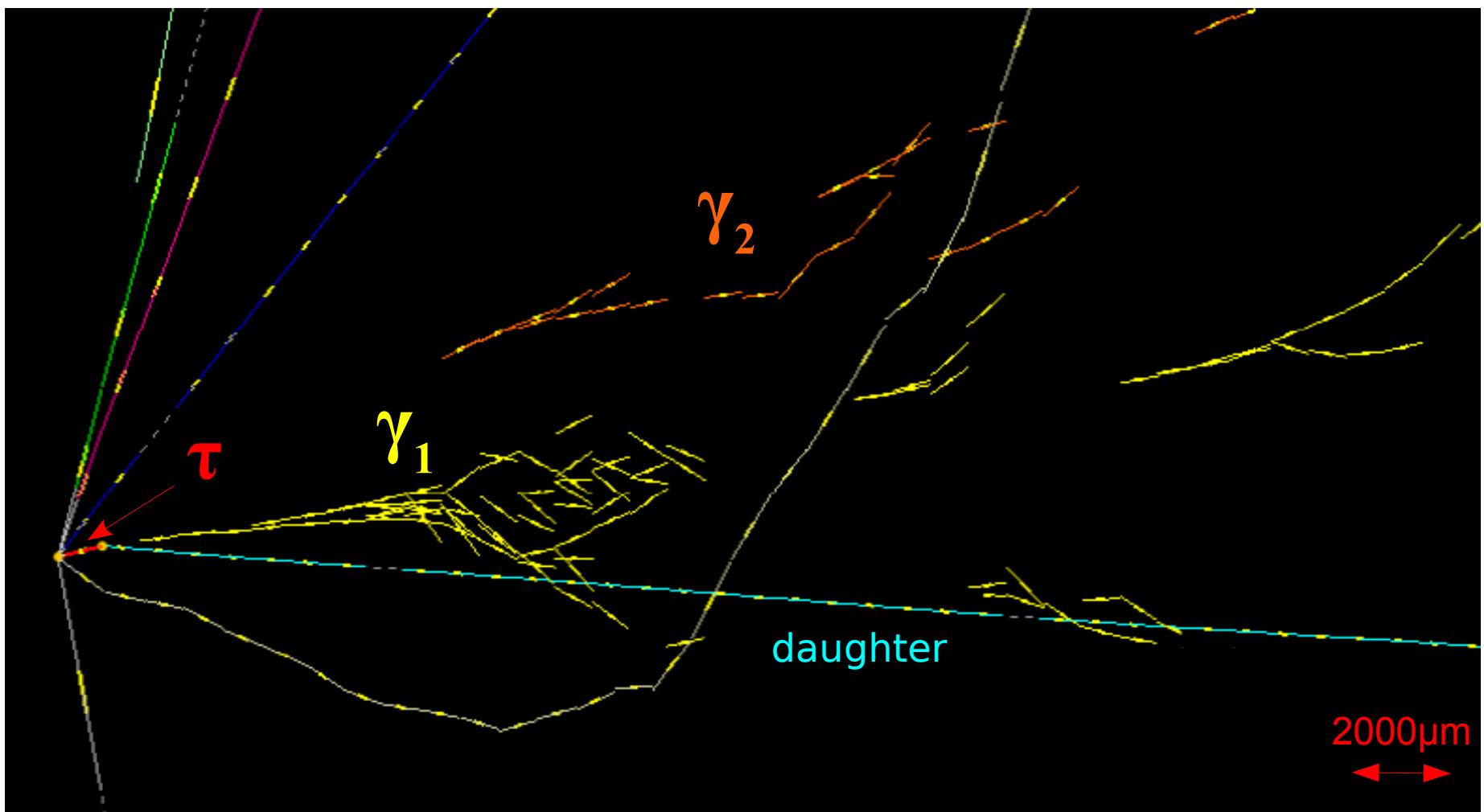
# search procedure status



# tau candidate event (1)



# tau candidate event (2)



# selection cuts, tau candidate

Variable	Cut-off	Value
Missing $P_T$ at primary vertex (GeV/c)	<1.0	$0.57^{+0.32}_{-0.17}$
Angle between parent track and primary hadronic shower in the transverse plane (rad)	$> \pi/2$	$3.01 \pm 0.03$
Kink angle (mrad)	$> 20$	$41 \pm 2$
Daughter momentum (GeV/c)	$> 2$	$12^{+6}_{-3}$
Daughter $P_T$ when $\gamma$ -ray at the decay vertex (GeV/c)	$> 0.3$	$0.47^{+0.24}_{-0.12}$
Decay length ( $\mu\text{m}$ )	<2 lead plates	$1335 \pm 35$

## ■ kinematical analysis:

- two EM showers ( $\gamma_1$  and  $\gamma_2$ ) pointing towards decay vertex,  
invariant mass:  $(120 \pm 20(\text{stat.}) \pm 35(\text{syst.}))\text{MeV}/c^2$   
**hypothesis:**  $\pi^0 \rightarrow \gamma\gamma$     ( $m_{\pi^0} = 135\text{MeV}/c^2$ )
- daughter is a charged hadron, most likely a charged pion,  
invariant mass ( $\pi + 2\gamma$ ):  $(640^{+125}_{-80} (\text{stat.})^{+100}_{-90} (\text{syst.}))\text{MeV}/c^2$   
**hypothesis:**  $\rho^- \rightarrow \pi^0 \pi^-$    ( $m_{\rho^-} = 770\text{MeV}/c^2$ )
- single-prong hadronic tau decay  
**hypothesis:**  $\tau^- \rightarrow \rho^- + \nu_\tau$    (B.R.  $\sim 25\%$ )  
$$\rho^- \rightarrow \pi^0 + \pi^-$$
  
$$\pi^0 \rightarrow \gamma\gamma$$

# tau background and efficiency

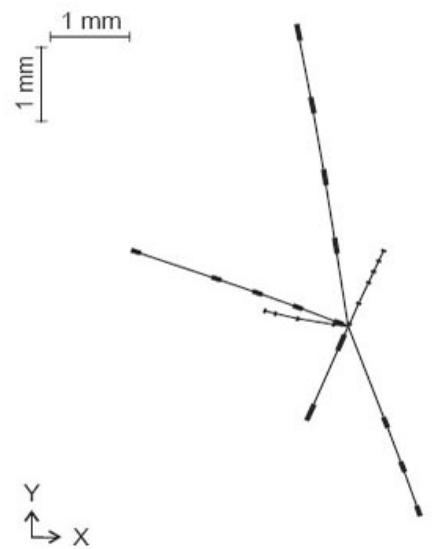
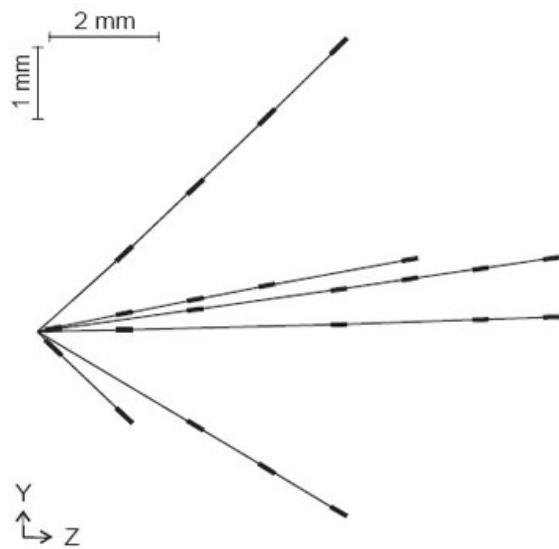
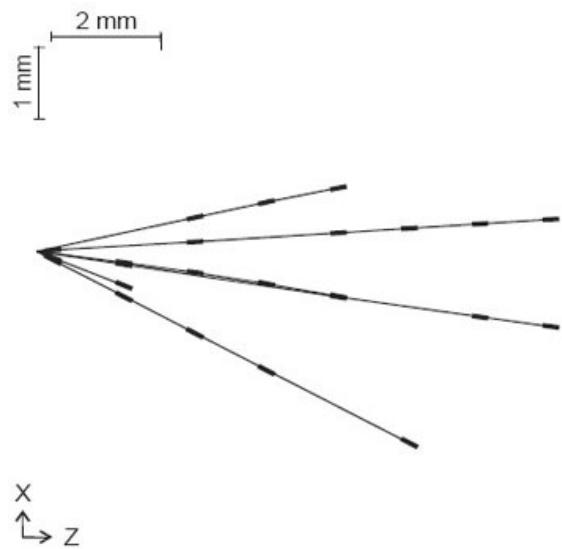
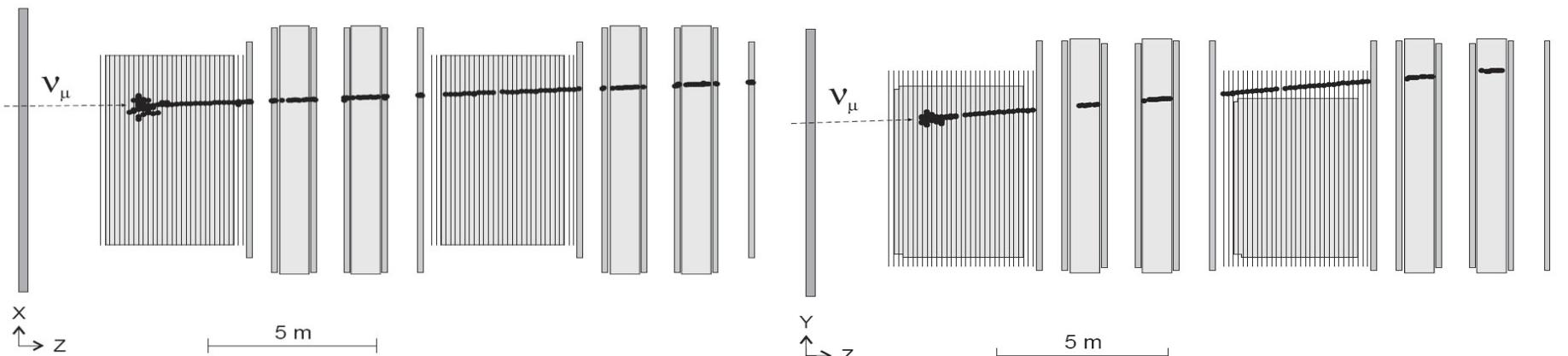
Decay channel	Number of background events expected for							
	$22.5 \times 10^{19}$ p.o.t.				$4.88 \times 10^{19}$ p.o.t.			
	Charm	Hadron	Muon	Total	Charm	Hadron	Muon	Total
$\tau \rightarrow \mu$	0.025	0.00	0.07	$0.09 \pm 0.04$	0.00	0.00	0.02	$0.02 \pm 0.01$
$\tau \rightarrow e$	0.22	0.00	0.00	$0.22 \pm 0.05$	0.05	0.00	0.00	$0.05 \pm 0.01$
$\tau \rightarrow h$	0.14	0.11	0.00	$0.24 \pm 0.06$	0.03	0.02	0.00	$0.05 \pm 0.01$
$\tau \rightarrow 3h$	0.18	0.00	0.00	$0.18 \pm 0.04$	0.04	0.00	0.00	$0.04 \pm 0.01$
Total	0.55	0.11	0.07	$0.73 \pm 0.15$	0.12	0.02	0.02	$0.16 \pm 0.03$

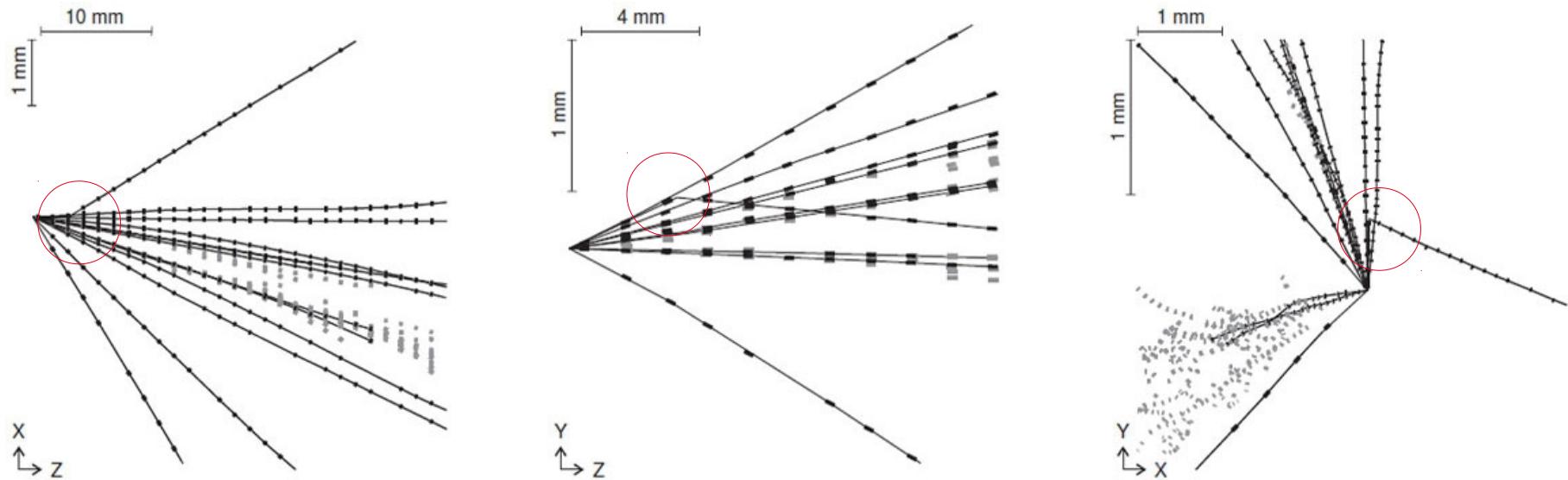
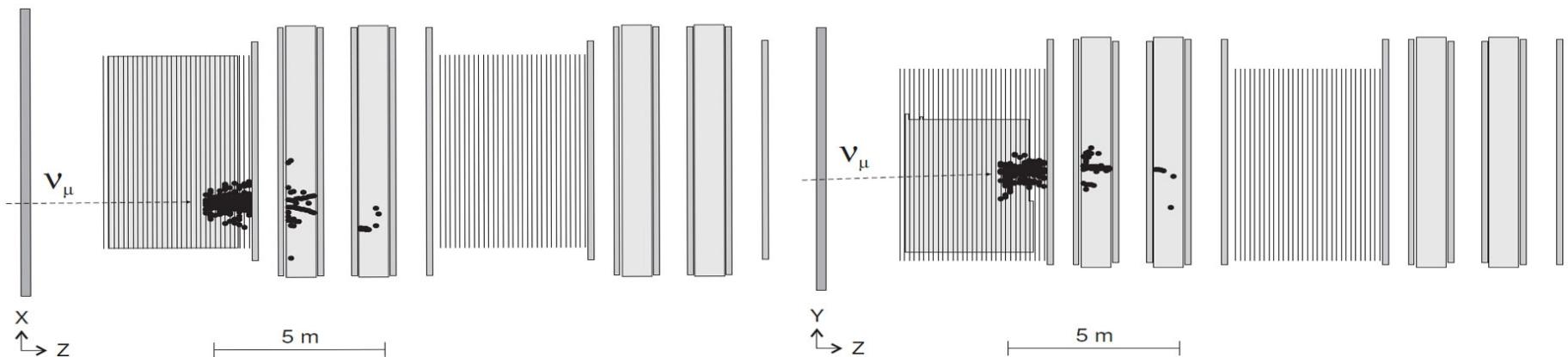
Decay channel	Number of signal events expected for		Interaction vertex location efficiency	Global $\tau$ detection efficiency
	$22.5 \times 10^{19}$ p.o.t.	$4.88 \times 10^{19}$ p.o.t.		
$\tau \rightarrow \mu$	1.79	0.39	0.54	0.09
$\tau \rightarrow e$	2.89	0.63	0.59	0.14
$\tau \rightarrow h$	2.25	0.49	0.59	0.04
$\tau \rightarrow 3h$	0.71	0.15	0.64	0.04
Total	7.63	1.65	0.59	0.07

# charm candidates

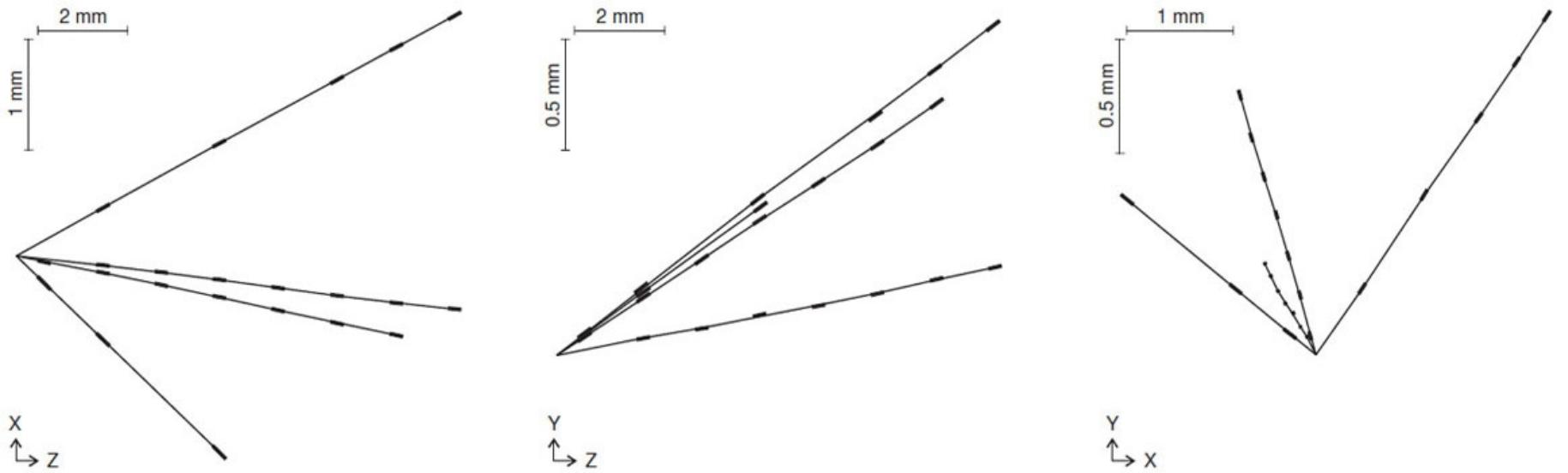
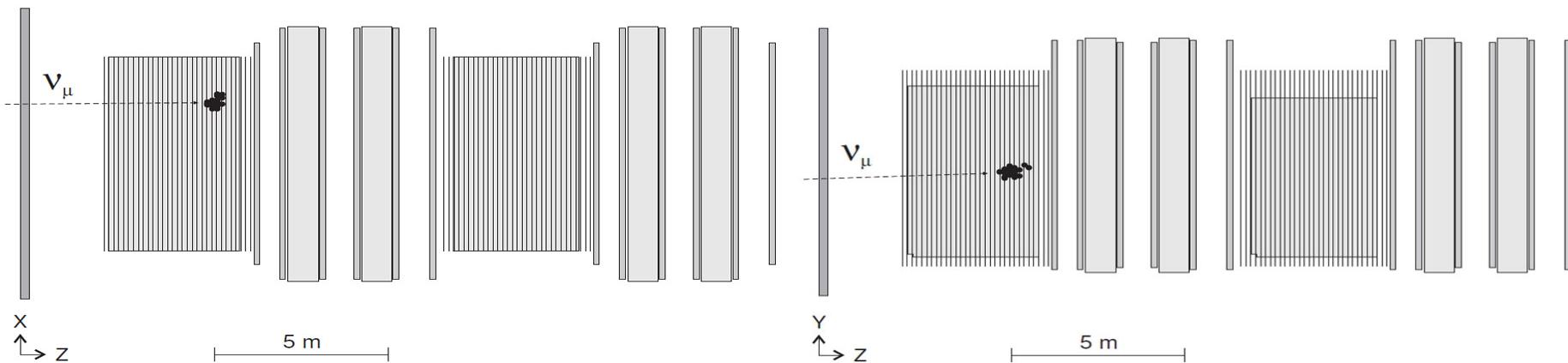
Topology	Observed charm candidate events	Expected events		
		Charm	Background	Total
Charged 1-prong	13	15.9	1.9	17.8
Neutral 2-prong	18	15.7	0.8	16.5
Charged 3-prong	5	5.5	0.3	5.8
Neutral 4-prong	3	2.0	<0.1	2.1
Total	39	$39.1 \pm 7.5$	$3.0 \pm 0.9$	$42.2 \pm 8.3$

# $\nu_\mu$ CC event

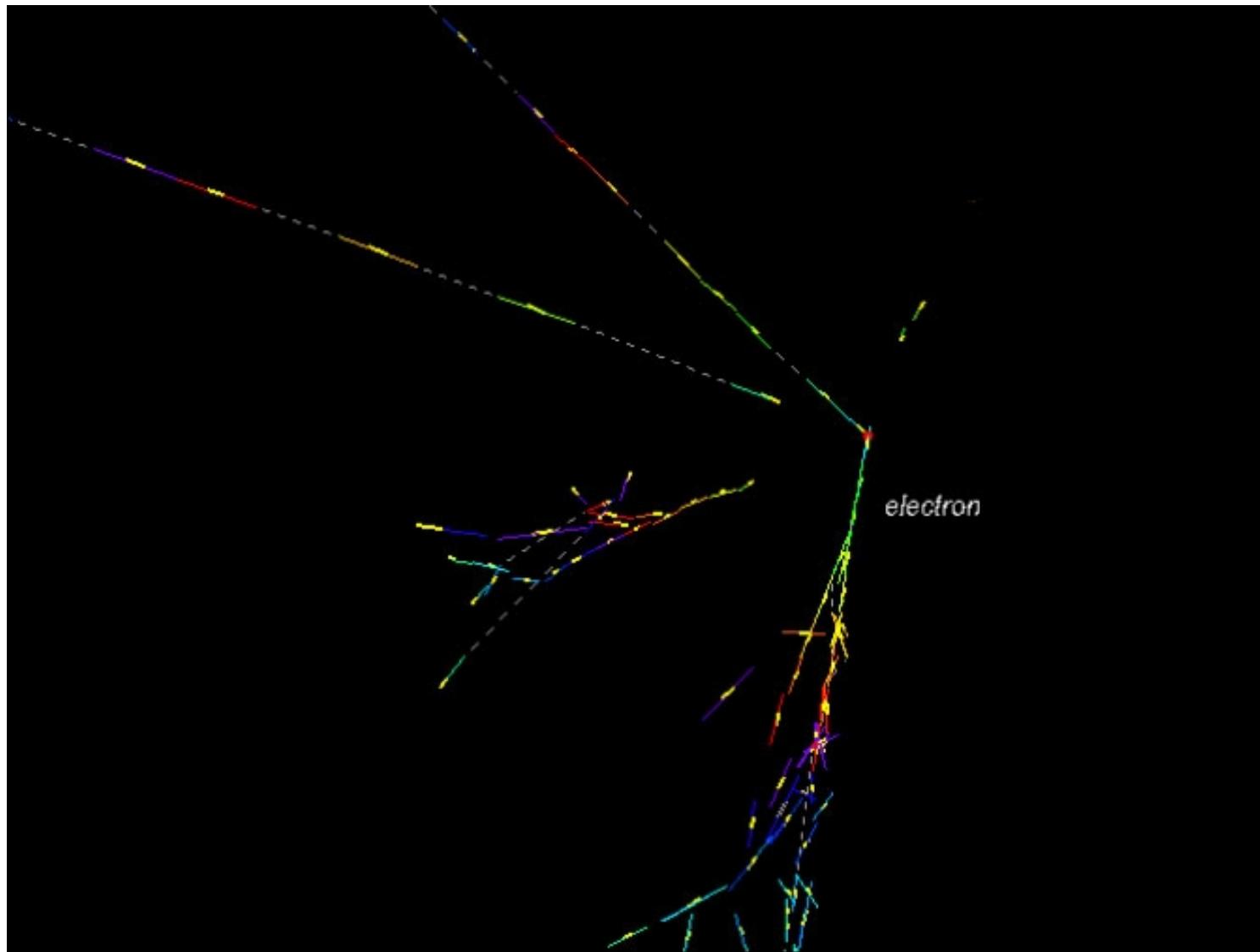


$\nu_\mu$  charm event

# $\nu$ NC event

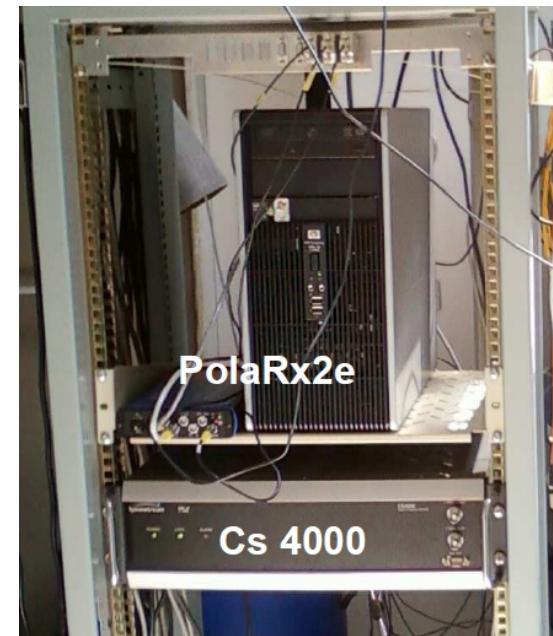


# $\nu_e$ candidate



# clock synchronization

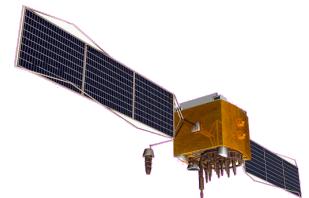
- identical system of GPS receivers and Cs clocks at CERN and LNGS
- use GPS “common view”: the **same satellites** seen by receivers at CERN and LNGS
- dual-frequency GPS,  
“ionosphere-free” P3 code
- locations at CERN and LNGS known with high accuracy
- calibrated by METAS,  
cross-checked by PTB
- establish “time-link”  $dt \sim 1\text{ns}$



# distance measurement

## ■ CERN ↔ LNGS above-ground: GPS

- establish new GPS benchmarks on both sides of the 10km highway tunnel
- measure reference GPS points at CERN and LNGS (2010)
- cross-checked CERN and LNGS reference points (June 2011)



## ■ LNGS ↔ OPERA underground: optical

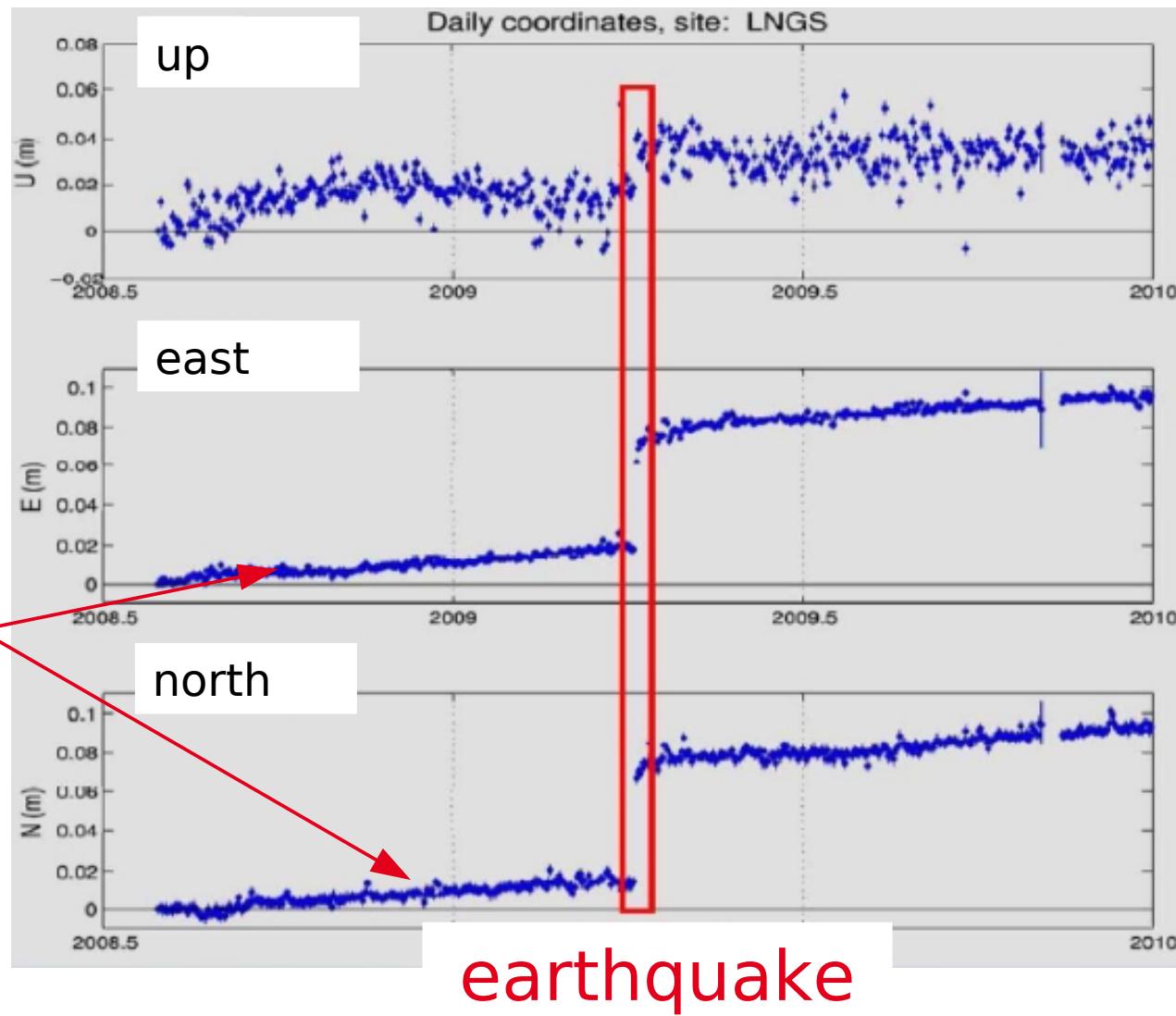
- block traffic (partially\*) on highway, use theodolites
- (\* reason for “bad” accuracy of only 0.2m)



$$d(\text{OPERA}_{\text{A}1} - \text{CERN}_{\text{BCT}}) = (731278.0 \pm 0.2) \text{ m}$$

# LNGS position monitoring

continental drift

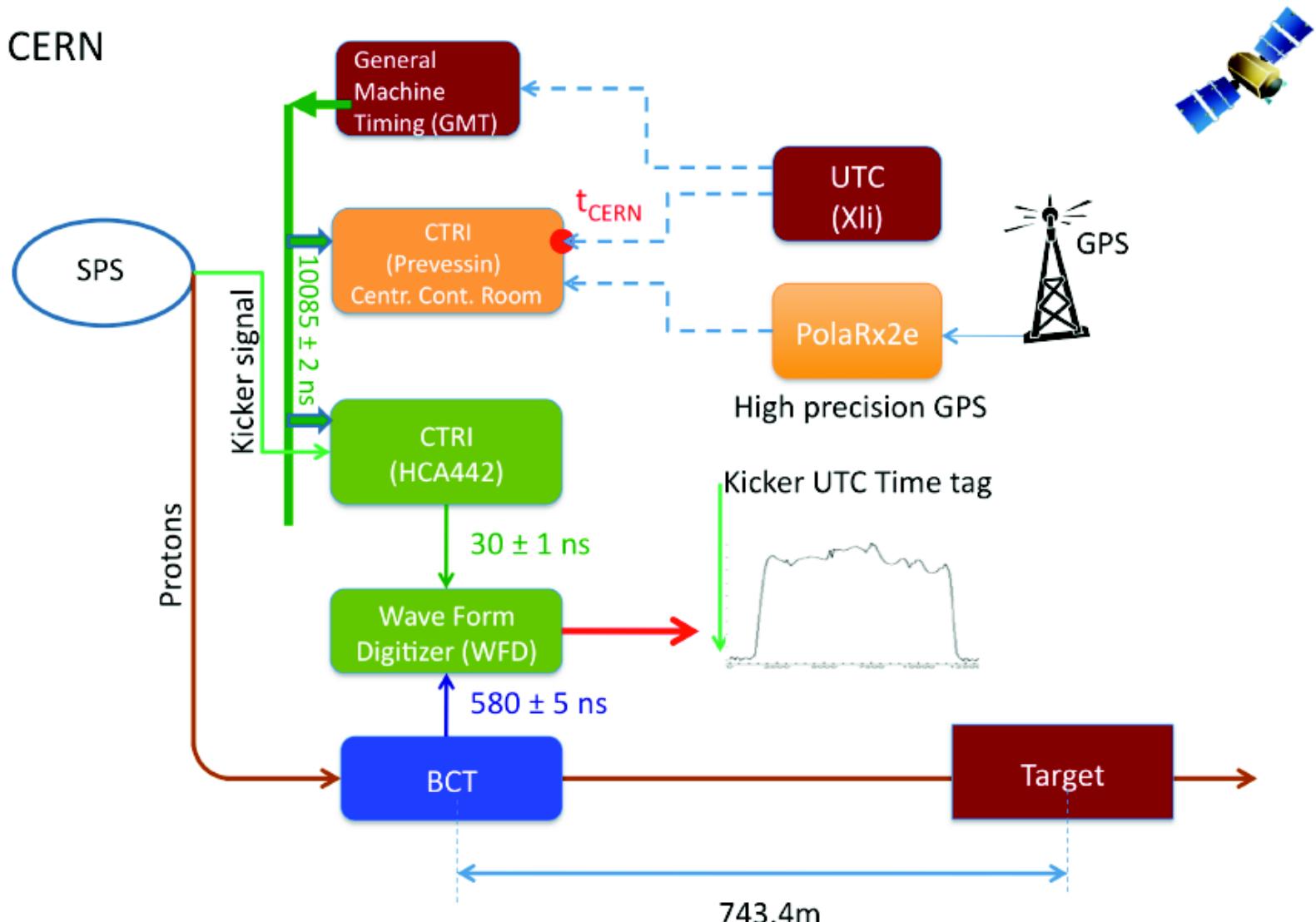


# systematic uncertainties

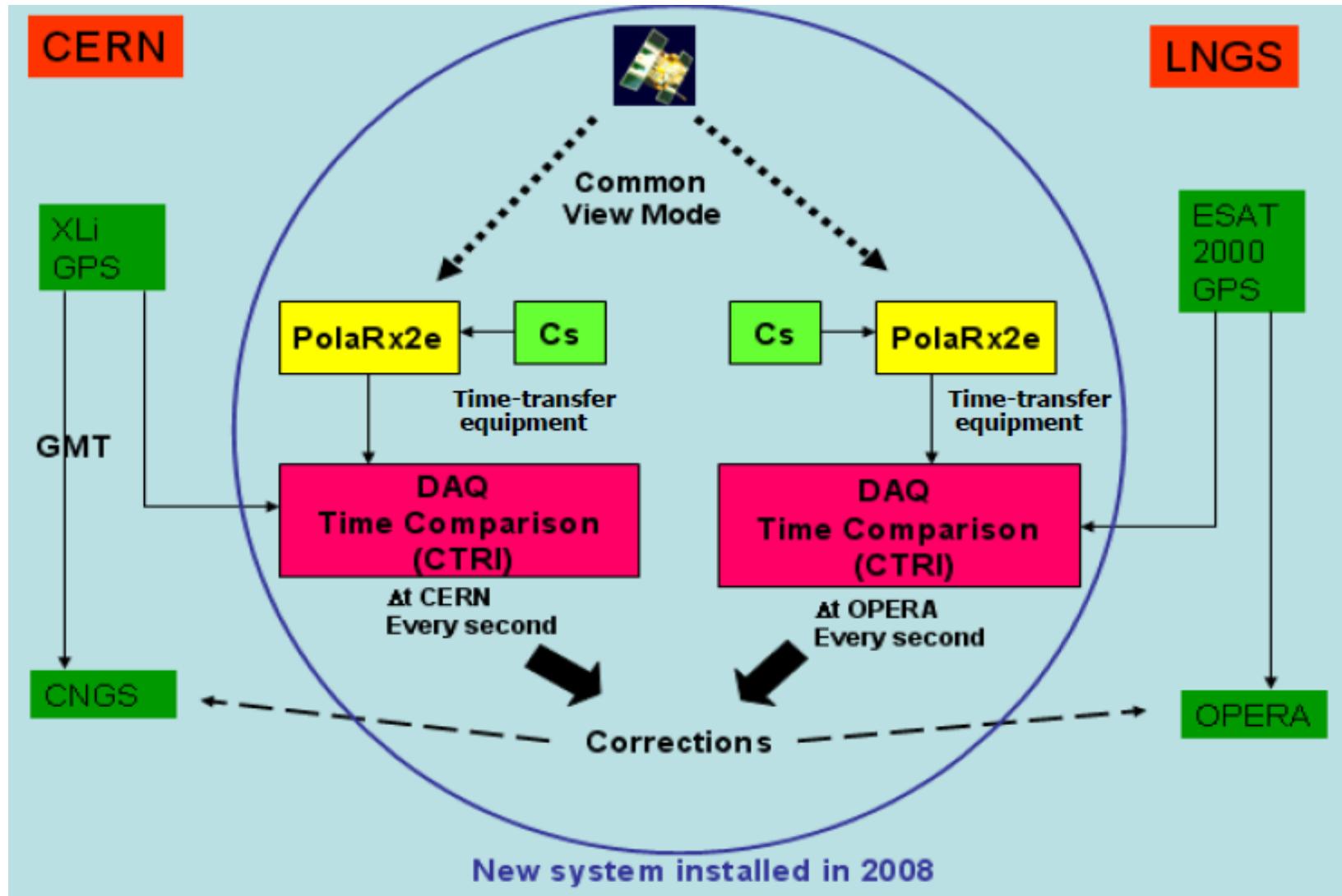
Systematic uncertainties	ns	Error distribution
Baseline (20 cm)	0.67	Gaussian
Decay point	0.2	Exponential (1 side)
Interaction point	2.0	Flat (1 side)
UTC delay	2.0	Gaussian
LNGS fibres	1.0	Gaussian
DAQ clock transmission	1.0	Gaussian
FPGA calibration	1.0	Gaussian
FWD trigger delay	1.0	Gaussian
CNGS-OPERA GPS synchronisation	1.7	Gaussian
MC simulation for TT timing	3.0	Gaussian
TT time response	2.3	Gaussian
BCT calibration	5.0	Gaussian
<b>Total systematic uncertainty</b>	<b>-5.9, +8.3</b>	

# CERN timing

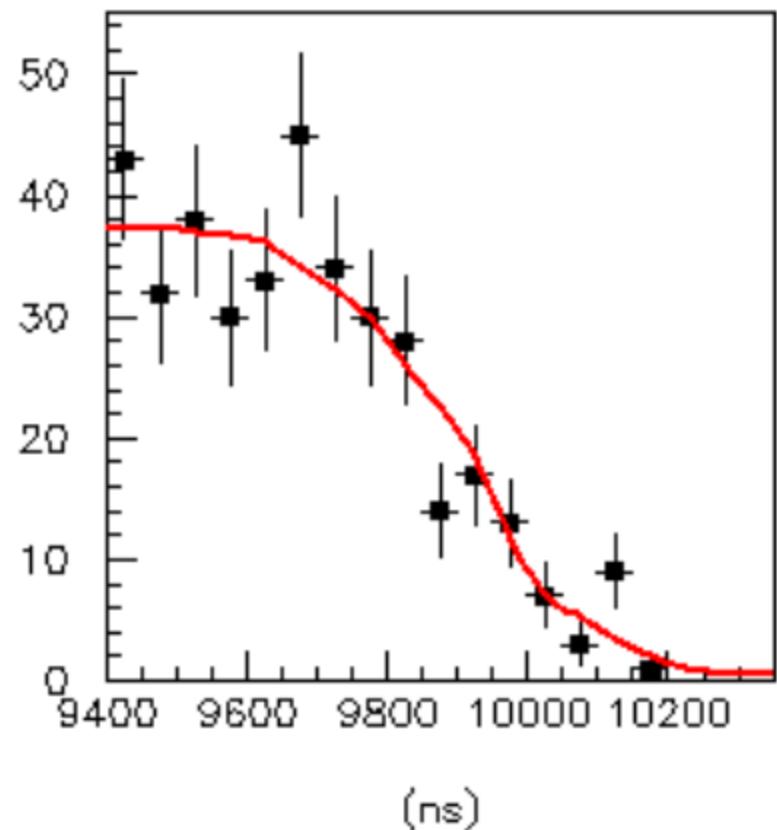
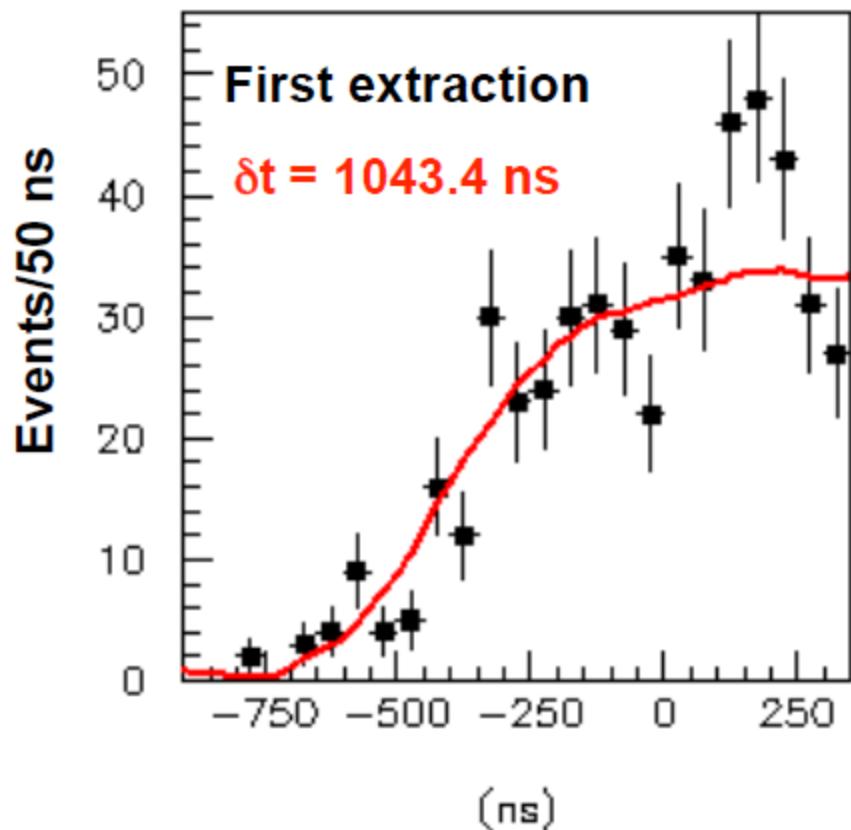
CERN



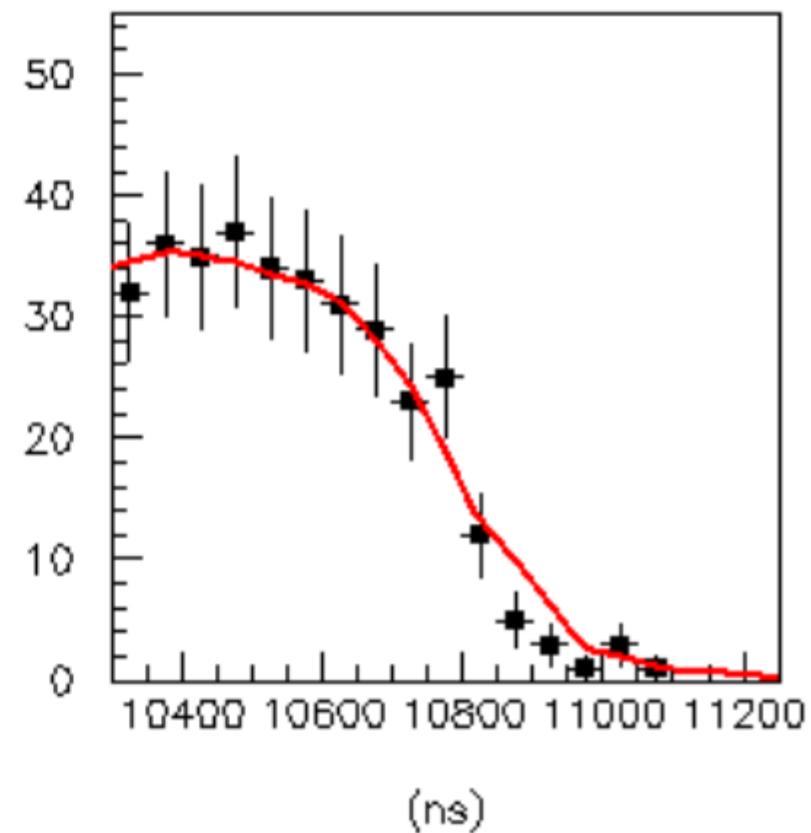
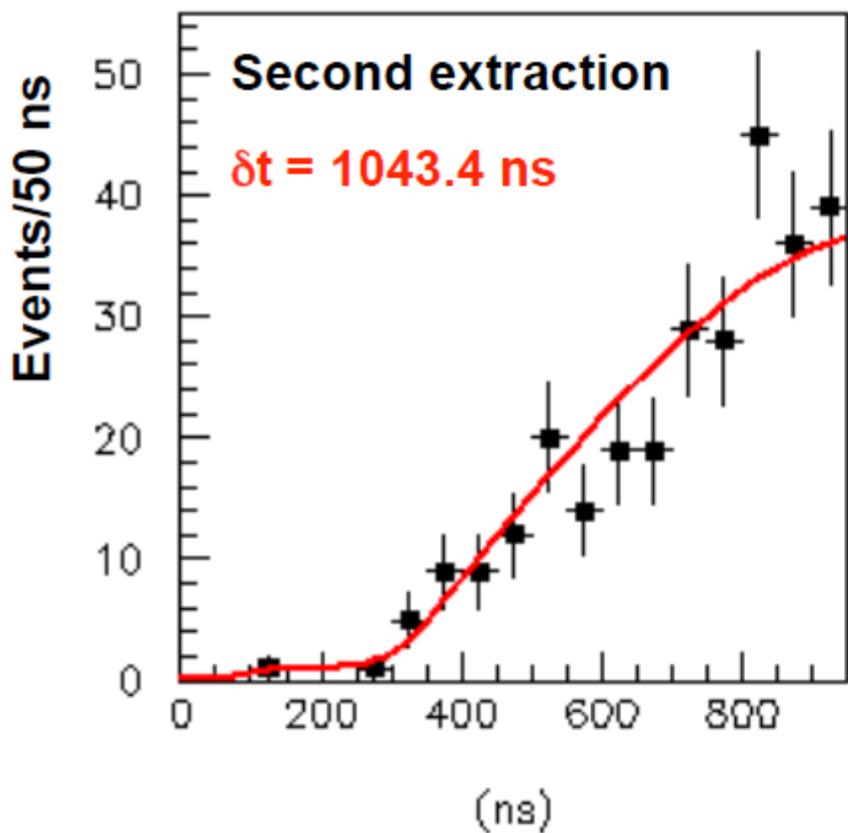
# clock synchronization

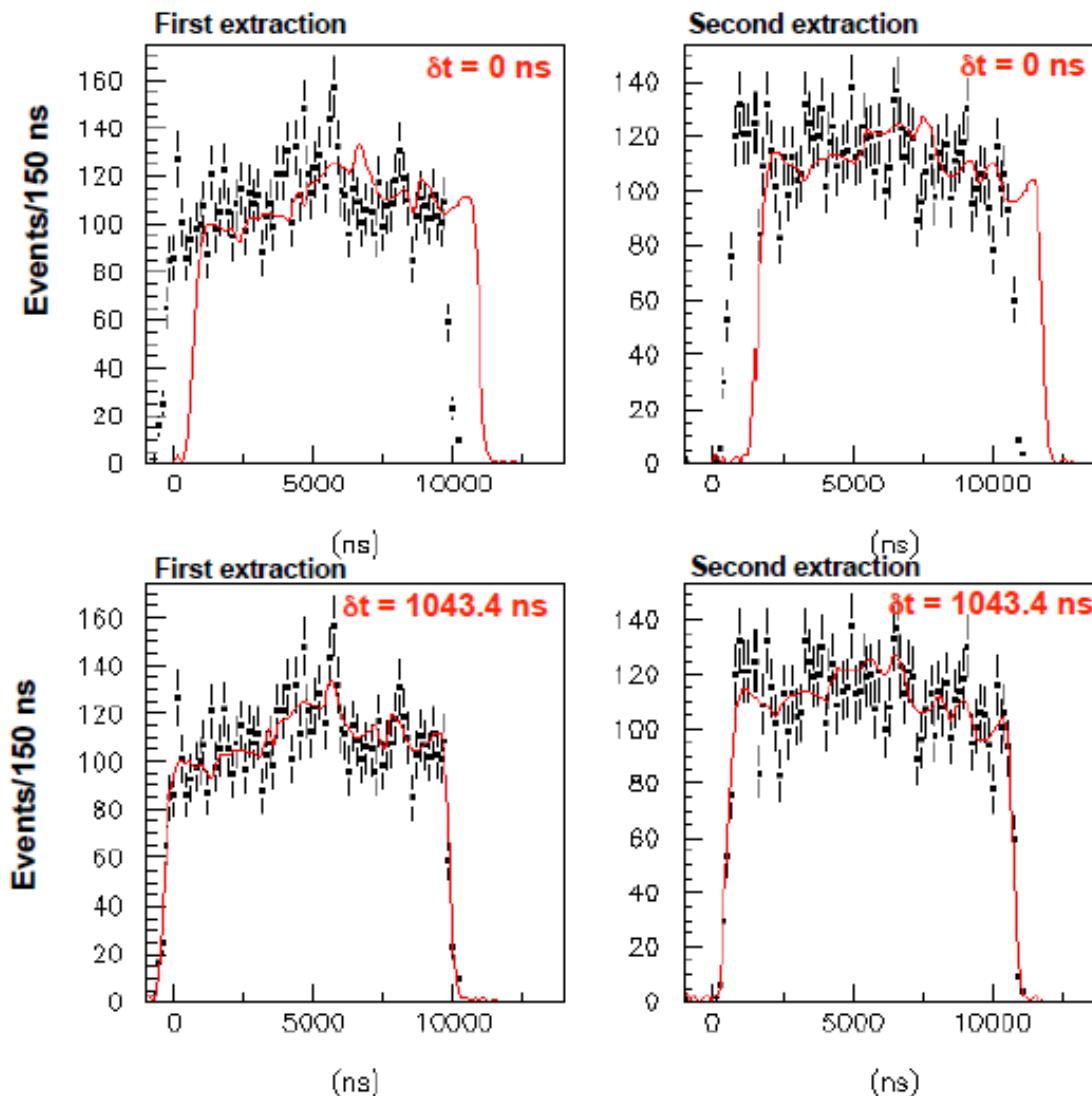


# edges extraction 1



# edges extraction 2





# first hit distribution

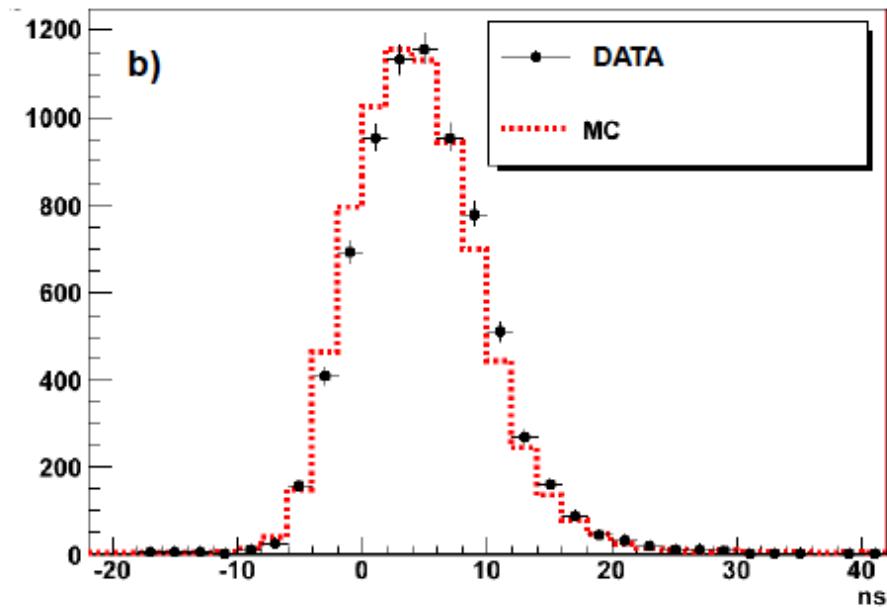
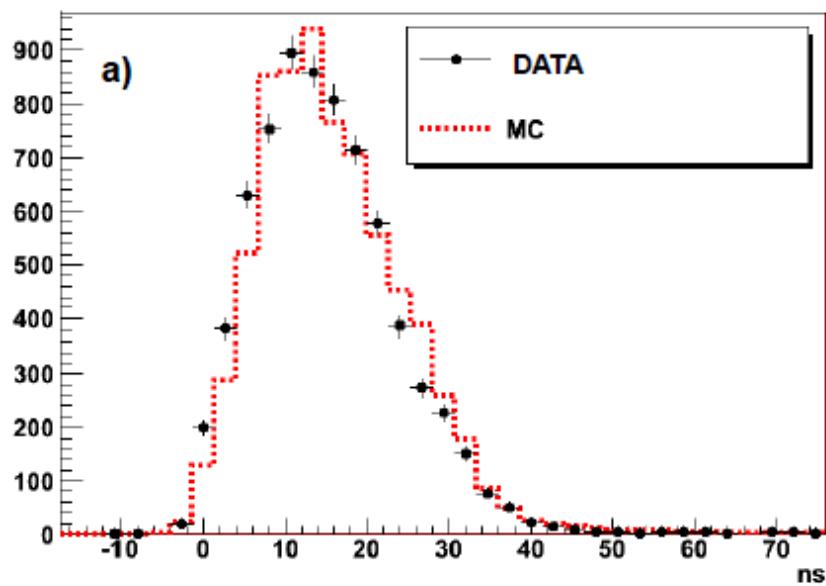


Fig. 9: Distribution of the time difference between the earliest TT hit and: a) the average time of the event, b) the average time of the muon track. Dots with error bars indicate data and the dotted line simulated events. Plot a) includes only internal events while plot b) only external events. The distributions are corrected for the longitudinal position of the hits.